

$$e^a{}_2 \rightarrow \lambda^a{}_b e^b{}_2$$

$$\Rightarrow e \rightarrow \lambda e$$

$$\text{Now } E = e^{-1} \rightarrow e^{-1} \lambda^{-1} \\ = \tilde{\lambda}^{-1}$$

$$\therefore E^M_a \xrightarrow[\text{local Lorentz trs.}]{} E^M_b (\tilde{\lambda}^{-1})^b_a$$

Suppose we have a ~~vector~~ field $A_\mu(x)$

$$\text{Define: } \hat{A}_a = E^M_a(x) A_\mu(x)$$

(we will now treat this as the indep. variable instead of $A_\mu(x)$)

We could take \hat{A}_a as indep. variables.

~~#~~ \hat{A}_a is scalar under general coordinate trs.

$$\hat{A}_a = E^M_a(x) A_\mu(x) \xrightarrow[\text{local Lorentz trs.}]{} E^M_b(x) A_\mu(x) (\tilde{\lambda}^{-1})^b_a \\ = \hat{A}_b(x) (\tilde{\lambda}^{-1})^b_a$$

$\therefore \hat{A}_a$ transforms as a covariant vector under local Lorentz trs.

Given A^M , we define:-

$$\hat{A}^a = \epsilon^a_{\mu} A^\mu$$

→ scalar under general
coord. & contravariant under local Lorentz trs.

$$\hat{A}^a(y) \rightarrow \Lambda_{\mu\nu}^{ab}(y) \hat{A}^b(y)$$

In general :-

Given $A^{\mu_1 \dots \mu_n}$ z_1, \dots, z_m

we define :-

$$A^{a_1 \dots a_n} {}_{b_1 \dots b_m} = \ell_{\mu_1}^{a_1} \ell_{\mu_2}^{a_2} \dots \ell_{\mu_n}^{a_n} \\ \times E^{v_1}_{b_1} \dots E^{v_m}_{b_m} \\ \times A^{\mu_1 \dots \mu_n} {}_{z_1 \dots z_m}$$

Guaranteed ?

whatever new action we get in terms of the vierbein should be inv. under local L.T.s. bcs the original action was inv. under local L.T.s.
— though this fact isn't manifest.

Possible confusion comes from derivatives
→ so what are the prop. which makes local L.T.s?

Q) How to define covariant derivatives
of tensors of local Lorentz trs. ?

$$\rightarrow D_\mu \hat{A}_\nu = E^v_b D_\mu A_v$$

Here there is no fundamental gauge field to compensate for the Do term, but here we can introduce indep. gauge fields for local L.T. — that will be too many d.o.f. So the normal way of defining covariant deriv. doesn't work here

invariant under local L.T. bcs A_v has no knowledge about local L.T., need to work this out

$$= E^v_b (\partial_\mu A_v - \Gamma^\rho_{\mu v} A_\rho)$$

$$\text{Now } A_\nu = e_2^a \hat{A}_a$$

We get

$$D_\mu \hat{A}_b = \tilde{\epsilon}_b^a \left\{ \partial_\mu (e_2^a \hat{A}_a) - \Gamma_{\mu\nu}^f e_2^a \hat{A}_f \right\}$$

Note :-

$\tilde{\epsilon}_b^a D_\mu A^\nu$
has been used
that gives
bcos $D_\mu \hat{A}_b$ the right
tr.s. prop. - we need
to decide what \hat{A}_b
we want $D_\mu \hat{A}_b$ to
be

$$e_2^a \partial_\mu \hat{A}_a + (\partial_\mu e_2^a) \hat{A}_a - \Gamma_{\mu\nu}^f e_2^a \hat{A}_f$$

$$= D_\mu \hat{A}_b + \hat{\omega}_{\mu b}^a \hat{A}_a$$

where

~~$$\hat{\omega}_{\mu b}^a = \tilde{\epsilon}_b^a \partial_\mu e_2^a - \tilde{\epsilon}_b^a \Gamma_{\mu\nu}^f e_2^f$$~~

$$\hat{\omega}_{\mu b}^a = \tilde{\epsilon}_b^a \partial_\mu e_2^a - \tilde{\epsilon}_b^a \Gamma_{\mu\nu}^f e_2^f$$

Or it's as if $\hat{\omega}_{\mu b}^a$ is a gauge field

- Instead of a gauge field here we have a complicated field det. in terms of e_2^a)

can be written in terms of
 e_μ^a & its inverse.

$$\text{Similarly, } D_\mu \hat{B}^b = e_2^b D_\mu B^2$$

$$= D_\mu \hat{B}^b + \hat{\omega}_{\mu b}^a \hat{B}^a$$

~~$$S.T. \hat{\omega}_{\mu b}^a = \tilde{\epsilon}_b^a \partial_\mu E^2 + \Gamma_{2f}^\mu \tilde{\epsilon}_f^a E^2$$~~

ω isn't an indep. field — it's det. in terms
of e^a_m 's

Define: $\hat{\omega}_\mu^{ba} = \eta^{bc} \hat{\omega}_{\mu c}^a$

$$\hat{\omega}_\mu^{ba} = \omega_\mu^b e_c^a \eta^{ca}$$

~~Ex:~~ Check that ① $\omega_\mu^{ab} = \hat{\omega}_\mu^{ab}$

② $\omega_\mu^{ab} = -\hat{\omega}_\mu^{ba}$

(to prove this, write P_m^a in terms e^a_m)

Bang! As long as we keep our connection fixed, we
don't have to use ~~ω~~ ω & $\hat{\omega}$ — can just
write ω (be careful about raising & lowering indices)

~~Ex:~~ S.T. $D_\mu(\hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_m})$

$$= D_\mu \hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_m} + \omega_\mu^{a_1} {}_{c_1} \hat{A}^{c_1 a_2 \dots a_n} {}_{b_1 \dots b_m} + \dots + \omega_\mu^{a_n} {}_{c_n} \hat{A}^{a_1 \dots a_{n-1} c_n} {}_{b_1 \dots b_m} + \omega_\mu^{d_1} {}_{b_1} \hat{A}^{a_1 \dots a_n} {}_{d_1 b_2 \dots b_m} + \dots + \omega_\mu^{d_m} {}_{b_m} \hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_{m-1} d_m}$$

$$D_\mu (\hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_m} \hat{B}^{c_1 \dots c_k} {}_{d_1 \dots d_l})$$

$$= D_\mu \left(\hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_m} \right)$$

$$\hat{B}^{c_1 \dots c_k} {}_{d_1 \dots d_l}$$

$$+ \hat{A}^{a_1 \dots a_n} {}_{b_1 \dots b_m} D_\mu \left(\hat{B}^{c_1 \dots c_k} {}_{d_1 \dots d_l} \right)$$

But why should these trs. as tensors of $SO(3, 1)$?

→ This would have been obvious had the $\omega_\mu{}^a{}_b$'s been gauge fields — but they aren't

$$\# \omega_\mu{}^b{}_a = e_2^b \partial_\mu E^a + T^\mu_{\nu\rho} e_2^b E^\rho_a$$

Under local Lorentz trs.,

$$e_2^b \xrightarrow[\text{to}]{\text{goes}} \Lambda_d^b e_2^d$$

$$E^\nu_a \longrightarrow \cancel{E^\nu_c} (\Lambda^{-1})^c{}_a$$

using these, we get

$$\omega_\mu{}^b{}_a \xrightarrow[\text{to}]{\text{goes}} \Lambda_d^b e_2^d \partial_\mu (E^c_c (\Lambda^{-1})^c{}_a)$$

$$+ T^\mu_{\nu\rho} \Lambda_d^b e_2^d E^\rho_c (\Lambda^{-1})^c{}_a$$

$$\Rightarrow \omega_\mu{}^b{}_a \rightarrow \Lambda_d^b \omega_\mu{}^d{}_c (\Lambda^{-1})^c{}_a$$

$$\text{Ex show this} \rightarrow + \Lambda_d^b \partial_\mu (\Lambda^{-1})^c{}_a$$

Define :- $\omega_\mu \Rightarrow$ matrix with ac component $\omega_{\mu a}^c$

(Then the eqn) $\omega_\mu \rightarrow \Lambda \omega_\mu \Lambda^{-1} + \Lambda \partial_\mu (\Lambda^{-1})$
 \Leftrightarrow exactly the way a Non-abelian gauge field transforms)

Here we didn't have the freedom to introduce new gauge fields — Nevertheless ω_μ transforms as if it is a gauge field

$$D_\mu \hat{A}^a = \partial_\mu \hat{A}^a + \omega_\mu^b \hat{A}^b$$

$$\hat{A}^a \rightarrow \Lambda^a_b \hat{A}^b$$

As a consequence $D_\mu \hat{A}^a \rightarrow \Lambda^a_b D_\mu \hat{A}^b \rightarrow$ transforms covariantly

$$\gamma^\mu = E^M_a \gamma^a$$

we'll now think this as fixed matrices

$$\{\gamma^a, \gamma^b\} = \eta^{ab}$$

Previously we had the problem that γ^μ are fixed matrices under general coord. trs.

The only way to define fermions is to erect a locally inertial frame & γ^i 's on it — can't define γ^i in a general coord. (or as a tensor under general coord. trs.) or as a tensor in a manifold.

Rep. of Lorentz grp

$dx^\mu \rightarrow$ Lorentz grp tensor

or ~~not~~ general coord. tensor

But Lorentz grp has tensors which aren't tensors under g.c.t. (can't be written as products of dx^μ)

→ So better write everything in terms of repn. of Lorentz grp.

It's an accident that some of the repn. of the Lorentz grp. can also be

~~7/11/08~~ $\omega_\mu{}^a{}_b = e_2^a \partial_\mu E_b{}^c + \Gamma_{\mu b}^{ac} e_2^a E_b{}^c$

Under local Lorentz ts.:-

$$e_2^a \rightarrow \lambda_b^a(x) e_2^b(x)$$

$$\omega_\mu{}^a{}_b \rightarrow \omega'_\mu{}^a{}_b = (\Lambda \partial_\mu \Lambda^{-1})^a{}_b + (\Lambda \omega_\mu \Lambda^{-1})^a{}_b$$

[It's like a non-abelian gauge ts. with gauge grp. $SO(3, 1)$]

$\omega_\mu{}^a{}_b$ transforms as a covariant vector under general coordinate transformation:-

$$\omega'_\mu{}^a{}_b = e_2^a \nu \partial_\mu E_b{}^c$$

→ transforms as a cov. vector under general coord. ts.

[Note: a, b indices don't ts. under general coord. ts. → they ts. only under local

Lorentz ts.]

$$\partial_\mu B^a = \partial_\mu B^a + \omega_\mu^{ab} B^b \text{ etc.}$$

→ trs. covariantly under ~~both~~ general
coord. trs. & contravariantly under local
local trs.

(This formalism is needed to deal with fermions
— not necessary for tensors under g.c.t.)

$$L = \bar{\Psi} (-i\gamma^\mu \partial_\mu \Psi - m \Psi)$$

→ no manifest 'a' index
here — so we ask whether
it is general coordinate
trs. invariant)

If Ψ transforms as a scalar, we have

$$\Psi'(x') = \Psi(x)$$

and clearly the above expression cannot
be general-coord. invariant.

So replace γ^μ by $\gamma^a E_a^\mu$

$$L = \bar{\Psi} (-i\gamma^a E_a^\mu \partial_\mu \Psi - m \Psi)$$

$$\Rightarrow L = \bar{\Psi} (-i\gamma^a E_a^\mu \partial_\mu \Psi - m \Psi)$$

→ manifestly general coord. trs.
invariant

But what about local Lorentz trs.?

→ Under this, $E_a^\mu \rightarrow E_b^\mu (\Lambda^{-1})^b{}_a$
We should take Ψ as trs. a spinor
under local Lor. trs.

$$\Psi'(\underline{x}) = R(N) \Psi(\underline{x})$$

or $\Psi'_{\alpha'}(\underline{x}') = (R(N))_{\alpha \beta} \Psi_{\beta}(\underline{x})$

↓ ↓ ↓
Spinor indices

$$L \xrightarrow{\text{becomes}} \Psi^+(R(N))^+ \gamma^0 \left[i \gamma^a \epsilon^{\mu}_{\nu}(N)_a^b \right. \\ \times \partial_{\mu} (R(N) \Psi) \\ \left. - m R(N) \Psi \right]$$

This is not a normal
 Lorentz trs. that we do in
 flat space — here we are
 dealing with an internal trs.
 e.g. $B^a(x) \rightarrow \Lambda^a_i(x) B^i(x)$

We are trying to guess
 what the trs. of Ψ should be — a priori
 we have no knowledge about it

Local Lor. trs. acts on Ψ doesn't act on \underline{x}
 at all — it's like an internal
sym. trs., where the relevant grp. is
 the Lorentz grp.

Local Lorentz trs. is only changing the
choice of vierbein & not the
coordinates.

$R(N)$
 ↓
 matrix
 acting
 on
 the
 spinor
 space
 for Lor.
 trs. matrix
 N

Under Lorentz trs.,
 $\int d^4x L(\Psi(x)) = \int d^4x' L(\Psi(x))$
 under local Lorentz trs.,
 $x' = x$ as it is
 an internal trs.

$$\begin{aligned}
 \mathcal{L} &\rightarrow \bar{\psi}^+ (\gamma^\mu)^\dagger \gamma^0 \left(-i \gamma^a E_a^\mu (\Lambda^{-1})_a^\mu \partial_\mu (R(\Lambda)) \bar{\psi} \right. \\
 &\quad \left. - m R(\Lambda) \bar{\psi} \right) \\
 \text{term ①} &\leftarrow = \bar{\psi}^+ R(\Lambda)^\dagger \gamma^0 \left(-i \gamma^a E_a^\mu (\Lambda^{-1})_a^\mu (\partial_\mu R(\Lambda)) \bar{\psi} \right) \\
 \text{term ②} &\leftarrow + \bar{\psi}^+ (\gamma^\mu)^\dagger \gamma^0 \left(-i \gamma^a E_a^\mu (\Lambda^{-1})_a^\mu R(\Lambda) \partial_\mu \bar{\psi} \right. \\
 &\quad \left. - m \bar{\psi}^+ R(\Lambda)^\dagger \gamma^0 R(\Lambda) \bar{\psi} \right)
 \end{aligned}$$

(At this pt. we are just trying out our guess that $\Psi'(x) = R(x) \bar{\psi} \gamma_\beta \psi(x)$ — we may have to

change our guess to make \mathcal{L} local-Lorentz inv.)
 don't even know whether there is any trs. which will make \mathcal{L} local-Lor.-inv.

Term ③ :-

(Need to use the prop. of $R(\Lambda)$)

Consider ordinary Lorentz trs. :-

$$\Psi'(x') = R(\Lambda_0) \Psi(x)$$

\hookrightarrow (not x -dep. for an ordinary L.T.)

$m \bar{\psi} \psi$ is invariant under an ordinary Lorentz trs. & hence

$$m \bar{\Psi}'(x') \Psi'(x') = m \bar{\Psi}(x) \Psi(x)$$

$$\Rightarrow m \bar{\Psi}^+(x) R(\Lambda_0)^\dagger \gamma^0 R(\Lambda_0) \Psi(x)$$

$$= m \bar{\Psi}^+(x) \gamma^0 \Psi(x)$$

By demanding $R(\Lambda_0)^\dagger \gamma^0 R(\Lambda_0) = \gamma^0$, we determine $R(\Lambda_0)$

So we learn that

$$R(\lambda_0)^+ \gamma^0 R(\lambda_0) = \gamma^0 \quad \textcircled{A}$$

So at a given pt. x ,

$$R(\lambda(x))^+ \gamma^0 R(\lambda(x)) = \gamma^0$$

bcos at each pt., $R(\lambda(x))$ satisfies property \textcircled{A} .

$$\begin{aligned} \text{So, } -m \psi^+ R(\lambda(x))^+ \gamma^0 R(\lambda(x)) \psi \\ = -m F(x) \psi(x) \end{aligned}$$

Term $\textcircled{2}$:-

$$\psi^+ (R(\lambda))^+ \gamma^0 \cancel{\gamma^a (\lambda^{-1})^b} \psi \text{ (if } \gamma_a^b)$$

$\swarrow \quad \downarrow \quad \searrow$

insert $R(\lambda) R(\lambda)^{-1}$

$$\text{Now, } R(\lambda)^{-1} \gamma^a R(\lambda) = \lambda^a c \gamma^c$$

To check whether the above relation is true \Rightarrow
 $L = \bar{\psi} (-i \gamma^\mu \partial_\mu + m \gamma^0)$ is inv. under

ordinary L.P.

$$\begin{aligned} L'(x^1) &= \bar{\psi}'(x^1) (-i \gamma^\mu \partial_\mu' + m \gamma^0) \\ &\stackrel{\text{known}}{=} \bar{\psi}(x) (-i \gamma^\mu \partial_\mu + m \gamma^0) \\ &\text{under } \begin{cases} \psi'(x^1) = \lambda_0 \psi(x) \\ x'^\mu = (\lambda_0)^M g_{\mu\nu} x^\nu \end{cases} \end{aligned}$$

$$L(x) = \psi^+(x) \left\{ R(N_0) + \gamma^0 (-i\gamma^\mu) (\tilde{\Lambda}^{-1})^\nu_\mu \right. \\ \times R(N_0) \left. \right\} \frac{\partial}{\partial x^\nu} \psi(x)$$

$\frac{\partial}{\partial x^\mu} = \frac{\partial x^\nu}{\partial x^\mu} \frac{\partial}{\partial x^\nu}$
 $= (\tilde{\Lambda}^{-1})^\nu_\mu \frac{\partial}{\partial x^\nu}$

↓ this bracket should be
equal to $(-i\gamma^0 \gamma^2)$

This is a specific prop. of $R(N_0)$
irrespective of $\psi(x)$

So we must have

$$R(N_0) + \gamma^0 \gamma^a R(N_0) (\tilde{\Lambda}^{-1})^b_a = \gamma^0 \gamma^b$$

for any N_0 γ^a 's are fixed matrices for flat space

[so taking γ^a , we can define γ^μ 's for curved space by $\gamma^\mu = \gamma^a E^\mu_a$]

γ^a 's don't trs. under g.c.t. — γ^μ 's do transform

$$\text{So, term } \textcircled{2} = \psi^+ \gamma^b \partial_\mu \psi (-i E^\mu_b) \\ = \bar{\psi} \gamma^b \partial_\mu \psi (-i E^\mu_b)$$

Hence, to summarize, we have,

$$L \xrightarrow[\text{local trs.}]{} L + \psi^+ R(N)^+ \gamma^0 (-i \gamma^\mu) E^\mu_b (\tilde{\Lambda}^{-1})^b_a \\ \times \partial_\mu R(N) \psi$$

* extra term [this is term $\textcircled{1}$]

Simplify the extra term:-

$$4 + \bar{g}^a \gamma^0 (-i\gamma^a) E_{\alpha}^{\mu} (\Lambda^{-1})^b_a (\bar{R}^{\alpha} R^{\beta}) \partial_{\mu} R^{\beta} \bar{g}$$

insert ~~over~~ this
here

So,

$$\mathcal{L} \rightarrow \mathcal{L} + \bar{\psi} (-i\gamma^b E_{\alpha}^{\mu}) R(\Lambda)^{-1} \partial_{\mu} R(\Lambda) \psi$$

Q) What is wrong? Why the extra term?

→ We have not used a covariant derivative for the local cov. trs. which is like a gauge trs.

\mathcal{L} has the global cov. trs.; but when we try to make it local, the sym. is broken — need to modify \mathcal{L} to make it invariant

Need to replace $\partial_{\mu} \bar{\psi}$ by $D_{\mu} \bar{\psi}$.

Q) What is $D_{\mu} \bar{\psi}$?

$$(D_{\mu} \bar{\psi})_{\alpha} = \partial_{\mu} \bar{\psi}_{\alpha} + (\bar{R}_{\mu})^{\beta \alpha} \bar{\psi}_{\beta}$$

↙
Spinor representation
of \bar{R}_{μ}

Need to use
the spinor repr.
of \bar{R}_{μ} VECTOR
INDICES
→ $(\bar{R}_{\mu})^{\alpha \beta}$
Spinor
index

$$C_n / C_m = \rho_r \rho_m \lambda$$

$$\left(\frac{f_r}{f_m} \right)_{t=t_0} \lambda(t_0) = \left(\frac{f_r}{f_m} \right)_{t=t_{eq}} \lambda(t_{eq})$$

$$\Rightarrow N(t_{eq}) = \left(\frac{f_r}{f_m} \right)_{t_0} \lambda(t_0)$$

$$\therefore RHS = 4 \left(\frac{c_0}{C_m} \right)^{1/4} \left(\frac{f_r}{f_m} \right)_{t_0}^{1/4} \lambda(t_0)^{3/4}$$

$$C_m = f_m \lambda^3$$

$$(C_m)^{-1/4} = (f_m)^{-1/4} \lambda^{-3/4}$$

$$\frac{f_m}{f_r} = 2.6 \times 10^4$$

$$\therefore RHS = 4 \left(\frac{c_0}{f_m} \right)_{t_0}^{1/4} \left(\frac{f_r}{f_m} \right)_{t_0}^{1/4}$$

$$= 4 \times \frac{10^{28}}{(2.6)^{1/4} \times 10} \times \frac{1}{3 \times (2.6)^{1/4} \times 10}$$

$$\simeq 10^{26}$$

$$\left. \begin{aligned} c_0^{1/4} &= 10^{15} \text{ GeV} \\ &= 10^{28} \text{ K} \end{aligned} \right\}$$

$$\left. \begin{aligned} \frac{f_m}{f_r} &= 2.6 \times 10^4 \\ f_m &= f_r \times 2.6 \times 10^4 \end{aligned} \right\}$$

$$f_r \propto T$$

$$f^N \gtrsim 10^{26}$$

$$\begin{aligned} N &\gtrsim 2.6 \ln 10 \\ &\gtrsim 6.0 \end{aligned}$$

$$g_2(t_1) = c_0 \quad \lambda(t_1) = \left(\frac{c_r}{c_0}\right)^{1/4}$$

End of com. const. dominated era:-

$$d(t_f) \approx \sqrt{\frac{3}{8\pi G c_0}} \frac{e^N}{\lambda(t_1)}$$

$$\lambda_{in} = \lambda(t_1)$$

$$g_{in} = c_0 \quad d(t_{ea}) = d(t_f) + \sqrt{\frac{3}{8\pi G c_0}} \lambda(t_{ea})$$

t : matter dominated era

$$d(t) = d(t_{ea}) + 2 \sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t)}$$

$$d(t_{rec}) = d(t_f) + 2 \sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t_{rec})} + \sqrt{\frac{3}{8\pi G c_R}} \lambda(t_{rec})$$

$$\lambda(t_{rec})$$

$$\sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t_{rec})}$$

$$C_r = C_m \lambda_{eq}$$

$$\lambda(t_1) = \left(\frac{C_m \lambda_{eq}}{C_r}\right)^{1/4}$$

$$\approx d(t_f) + 2 \sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t_{rec})} \quad \lambda(t_{ea}) < \lambda(t_{rec})$$

$$\Delta = 2 \left(d(t_b) - d(t_{rec}) \right) = 2 \sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t_b)/\lambda(t_b)} \gg \lambda(t_{rec})$$

We want $\frac{4}{d(t_{rec})} < 1$

$$\Rightarrow \Delta < d(t_{rec}) \quad \Rightarrow 4 \sqrt{\frac{3}{8\pi G c_m}} \sqrt{\lambda(t_b)} < \sqrt{\frac{3}{8\pi G c_0}} \frac{e^N}{\lambda(t_1)}$$

$$\Rightarrow e^N > 4 \sqrt{\frac{c_0}{c_m}} \lambda(t_1) \sqrt{\lambda(t_b)}$$

$$\text{RHS} = 4 \sqrt{\frac{c_0}{c_m}} \left(\frac{c_m}{c_0}\right)^{1/4} (\lambda_{eq})^{1/4} \sqrt{\lambda(t_b)}$$

$$= 4 (c_0/c_m)^{1/4} (\lambda_{eq})^{1/4} \sqrt{\lambda(t_b)}$$

Steps to get ∂_μ :

In non-abelian gauge theory, we have gauge fields A_μ^a .

Suppose T^a are the generators in some representation.

$$D_\mu \phi = \partial_\mu \phi + i A_\mu^a T^a \phi \text{ if } \phi \text{ is in repn'g}$$

Suppose X is a field in a different repn.

$$\text{then } D_\mu X = \partial_\mu X + i A_\mu^a R(T^a) X$$

\downarrow
generator in the repn for X

(Make sure generators are normalised in the same way in all the repns :)

$$\text{if } [T^a, T^b] = i f^{abc} T^c$$

$$\text{then } [R(T^a), R(T^b)] = i f^{abc} R(T^c)$$

Then if $D_\mu \phi$ transforms covariantly,
then $D_\mu X$ also "

" , No need to worry about $\text{Tr}[T^a T^b] = 0$

$$\begin{aligned} D_\mu B^a &= \partial_\mu B^a + \omega_\mu^{ab} B^b \\ &= \partial_\mu B^a + i \sum_{cd} \omega_\mu^{cd} (T^{cd})^a {}_b B^b \end{aligned}$$

where $(T^{cd})^a {}_b = \delta^a {}_c \eta_{bd} - \delta^a {}_d \eta_{bc}$

i. $\star \leftrightarrow_{\text{fw}} (\text{cd})$ with $c < d$,
 [we have 6 such pairs = ${}^4C_2 = \frac{4!}{2 \times 2}$]

~~\star~~ $\star \leftrightarrow_{\text{fw}} w_\mu^{cd}$,

and $(T^k)^a{}_b \leftrightarrow_{\text{fw}} (T_{cd})^a{}_b$

Now, $[T_{cd}, T_{ef}] = i f_{(cd)(ef)(gh)} T_{gh}$
 (ensuring $c < d, e < f, g < h$
 for each pair)

~~so~~ $(\Omega_\mu)_{\alpha\beta} = i w_\mu^{cd} \{ R(T_{cd}) \}_{\alpha\beta}$

L (\rightarrow prescription to define the cov. deriv.
 for T)

We know that

$[R(T_{cd}), R(T_{ef})] = i f_{(cd)(ef)(gh)} R(T_{gh})$

↓ Lorentz generators
 in the spinor & repn.

$$R(T_{cd}) = \star (\gamma_c \gamma_d - \gamma_d \gamma_c)$$

↓ Need to fix the normalisation 'K'
 by using the commutation

$$\text{relation } [R(T_{cd}), R(T_{eg})] = i f_{(cd)(eg)(gh)} R(T_{gh})$$

Ex/ Find K

$$D_\mu \Phi = \partial_\mu \Phi + i K \omega_\mu^{cd} (\epsilon_c \gamma_d - \epsilon_d \gamma_c) \Phi$$

So the lagrangian now is

$$\mathcal{L} = \bar{\Phi} \left(-i g \gamma^\mu D_\mu \Phi - m \Phi \right)$$

$\approx \gamma^a E^{\mu}_{a}$

Ex: Check explicitly that \mathcal{L} is local Lorentz invariant.

\mathcal{L} is also inv. under g.c.t. as ω_μ^{cd} is a vector under g.c.t. (so $E^\mu \omega_\mu^{cd}$ is a scalar) & γ^a 's are inv. under g.c.t.

Only diff. from Non-abelian gauge theory is that there we use an indep. \mathbf{B} and gauge field A_μ^k (an indep. entity), while here ω_μ^{ad} 's are constructed out of \dots — however it trs. ~~as if~~ as if it were a gauge field

If we fix T^a, T^b & take same Tr $[T^a, T^b] = \#$,

$R(T^a), R(T^b)$ will be det. by

$$[R(T^a), R(T^b)] = \text{if } a \neq b, R(T^c) - \textcircled{A}$$

& $\text{Tr}(R(T^a) R(T^b))$ will also be det.

by \textcircled{A} .

In $\text{Tr}(F_{\mu\nu} F^{\mu\nu})$ we need only $\text{Tr}(T^a T^b) \#$ which is fixed by our choice.

15

Generalisation of Schwarzschild

Black holes to Black holes carrying charge

Schwarzschild metric:-

$$ds^2 = -\gamma(f) dt^2 + \chi(f) df^2 + f^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

$$(y^1, y^2, y^3) = (f \sin \theta \cos \phi, f \sin \theta \sin \phi, f \cos \theta)$$

$$\begin{pmatrix} y^1 \\ y^2 \\ y^3 \end{pmatrix} \rightarrow R \begin{pmatrix} y^1 \\ y^2 \\ y^3 \end{pmatrix} \quad (\text{spherical sym.})$$

(Action of Maxwell's th. coupled to Einstein's) \rightarrow

$$S = -\frac{1}{16\pi G} \int \sqrt{-g} R d^4x$$

$$- \frac{1}{4} \int d^4x \sqrt{-g} F_{\mu\nu} F^{\mu\nu}$$

It's better not to try to figure out the form of A_μ to get the form of $F_{\mu\nu}$.
Because A_μ 's are gauge-dependent \Rightarrow so try to find an ansatz for A_μ itself.

$$\cancel{F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu}$$

$$F_{t\mu} = \tilde{f}(f) y^\mu$$

$$F_{ij} = \epsilon_{ijk} y^k \tilde{g}(f)$$

$$F_{t\mu} = \partial_\mu A_t - \partial_t A_\mu$$

$A_0 = f(s) \quad A_i = g(s)$
Won't work — Don't think it's form as far as far as overall form

beats y^i trs. as a vector under rot. \rightarrow so $F_{t\mu}$ will have the same prop.

no time dep.
 \rightarrow so far of 'f' only

Ex: Show that in t, f, θ, ϕ coordinates

$$F_{t\mu} = f(f)$$

$$f(f) = f \tilde{f}(f)$$

$$F_{\theta\phi} = g(f) \sin \theta$$

All other components
 0

We now need to determine $\gamma(f)$, $\chi(f)$, $f(f)$ and $g(f)$.

We have :-

- ① Maxwell's eqns.
- ② Einstein's eqns.
- ③ Bianchi Identity

In flat space

$$T_{\text{e.m.}}^{\mu\nu} = \eta^{rs} \eta^{\mu\sigma} \eta^{\nu\tau} F_{rK} F_{\sigma\tau} - \frac{1}{4} \eta^{\mu\nu} \eta^{rs} \eta^{\nu\tau} F_{rK} F_{\sigma\tau}$$

$$T_{\text{matter}}^{\mu\nu} = \sum_m \int d^4x \delta^{(u)}(x^i - x_m^i) \frac{dx^a}{dx} \frac{dx^m}{dx}$$

$$\begin{aligned} T_{\text{em}}^{00} &= g^{0s} g^{0c} g^{KT} F_{rK} F_{\sigma\tau} - \frac{1}{4} g^{00} g^{00} g^{KT} \\ &\quad \times F_{rK} F_{\sigma\tau} \\ &= g^{00} g^{00} g^{KT} F_{0K} F_{0\tau} + \frac{1}{4} \left[g^{00} g^{KT} F_{0K} F_{0\tau} \right. \\ &\quad \left. + g^{ii} g^{KT} F_{ik} F_{it} \right] \\ &= \frac{1}{4} g^{ss} (F_{0s})^2 + \frac{1}{4} \left[-\frac{1}{4} g^{ii} (F_{0i})^2 \right. \\ &\quad \left. + g^{ii} g^{00} (F_{0i})^2 \right. \\ &\quad \left. + g^{ii} g^{kk} F_{ik} F_{ik} \right] \\ &= \frac{1}{4} g^{ss} (F_{0s})^2 \\ &\quad + \frac{1}{4} \left[-\frac{1}{4} g^{ss} (F_{0s})^2 - \frac{g^{ss}}{4} (F_{0s})^2 \right. \\ &\quad \left. + g^{00} g^{00} F_{0\phi} F_{0\phi} \times 2 \right] \\ &= \frac{f^2}{4^2 x} + \frac{1}{4} \left[-\frac{f^2}{4x} - \frac{f^2}{4x} + \frac{2}{f^4 \sin^2 \theta} g^2 \sin^2 \theta \right] \end{aligned}$$

$$= \frac{f^2}{4^2 x} - \frac{f^2}{24^2 x} + \frac{g^2}{24 f^4}$$

Can't use ~~easy~~ div. of $T_{\mu\nu}$ zero bcos that follows as a consequence of e.o.m.

bcoz it's not a priori guaranteed that the ansatz for $F_{\mu\nu}$ we have taken will satisfy Bianchi identity — of course ~~guarantees~~ would have guaranteed that Bianchi identity is satisfied

$$T_{em}^{00} = + \frac{f^2}{2+x^2} + \frac{g^2}{2+f^4}$$

$$T_{em}^{SS} = g^{SS} g^{S\bar{S}} g^{K\bar{K}} F_{PK} F_{SK}$$

$$- \frac{1}{4} g^{SS} g^{80} g^{K\bar{K}} F_{8K} F_{SK}$$

$$= \frac{1}{x^2} \left[g^{tt} F_{pt} F_{st} \right] - \frac{1}{4x} \left[g^{tt} g^{ss} F_{ts} F_{tp} \times 2 + g^{00} g^{ph} F_{0\phi} F_{0p} \times 2 \right]$$

$$= - \frac{1}{x^2} f^2 - \frac{1}{4x} \left[- \frac{2f^2}{4+x} + \frac{2}{f^4 \sin^2 \theta} \right]$$

$$= - \frac{f^2}{x^2} + \frac{f^2}{2+4x^2} - \frac{g^2}{2x f^4}$$

$$= - \frac{f^2}{2+x^2} - \frac{g^2}{2x f^4}$$

$$T_{em}^{ph} = g^{ph} g^{p\bar{0}} g^{K\bar{K}} F_{8K} F_{SK} - \frac{1}{4} g^{ph} g^{80} g^{K\bar{K}} F_{8K} F_{SK}$$

$$= g^{ph} g^{p\bar{0}} g^{00} F_{0\phi} F_{0\phi}$$

$$- \frac{1}{4} g^{ph} \left[- \frac{2f^2}{4x} + \frac{2g^2}{f^4} \right]$$

$$= \frac{g^2 \sin^2 \theta}{f^6 \sin^4 \theta} - \frac{1}{2f^2 \sin^2 \theta} \left[- \frac{2f^2}{4x} + \frac{g^2}{f^4} \right]$$

$$= \frac{g^2}{f^6 \sin^2 \theta} + \frac{f^2}{2f^2 \sin^2 \theta + x} - \frac{g^2}{2f^6 \sin^2 \theta}$$

$$= + \frac{g^2}{2f^6 \sin^2\theta} + \frac{f^2}{2f^2 \sin^2\theta + x}$$

$$\begin{aligned}
 T_{em}^{00} &= g^{00} g^{00} g^{\phi\phi} F_{0\phi} F_{0\phi} - \frac{1}{4} g^{00} [g^{\delta\delta} g^{kk} F_{sk} F_{sc}] \\
 &= \frac{g^2 \cancel{\sin^2\theta}}{f^6 \cancel{\sin^2\theta}} - \frac{1}{4 f^2} \left[-\frac{2f^2}{4+x} + \frac{2g^2}{f^4} \right] \\
 &= \frac{g^2}{f^6} + \frac{f^2}{2f^2+x} - \frac{g^2}{2f^6} \\
 &= \frac{g^2}{2f^6} + \frac{f^2}{2f^2+x}
 \end{aligned}$$

$$T_{em}^{oi} = 0$$

(fix)

$$\begin{aligned}
 T_{em}^{ij} &= g^{i\delta} g^{j\phi} g^{kk} F_{\delta k} F_{\phi c} \\
 &\quad - \cancel{\frac{1}{4} g^{ij} [- -]}
 \end{aligned}$$

$$= g^{ii} g^{jj} g^{kk} F_{ik} F_{jc}$$

$$\begin{aligned}
 &= g^{ii} g^{jj} \cancel{g^{00} F_{j0} F_{i0}} \\
 &\quad + g^{ii} g^{jj} \cancel{g^{ll} F_{il} F_{jl}}
 \end{aligned}$$

$$= 0$$

To solve :-

$$R_{\mu\nu} = 4\pi G \left[g^{\alpha\sigma} g_{\mu\nu} T_{\alpha\sigma} - 2 T_{\mu\nu} \right]$$

~~t t - component~~

~~$$R_{tt} = 4\pi G \left[g^{\alpha\sigma} g_{tt} T_{\alpha\sigma} - 2 T_{tt} \right]$$~~

$$\Rightarrow \frac{R_{tt}}{4\pi G} = \cancel{g_{tt} T_{tt}}$$

$$= g_{tt} \left[-\frac{f^2}{2x^3 f} - \frac{g^2}{f^4 f^2} - \frac{f^2}{2x^3 f} \right]$$

$$T_{tt} = \frac{1}{2f^2} \left[\frac{g^2}{f^4} + \frac{g^2}{f^4} \right] = \frac{g^2}{2f^4} + \frac{g^2}{2f^4}$$

~~$$T_{\theta\theta} = g^{\phi\phi} g_{\theta\theta} T_{\theta\theta}$$~~
$$= -\frac{1}{2x^2} \left[\frac{f^2}{x^2 f} + \frac{g^2}{x f^4} \right]$$
$$= -\frac{f^2}{2x^4 f} - \frac{g^2}{2x^3 f^4}$$

~~$$T_{\phi\phi} = \frac{1}{2f^2 \sin^2 \theta} \left[\frac{g^2}{f^6 \sin^2 \theta} + \frac{f^2}{f^2 \sin^2 \theta f^2} \right]$$~~
$$= \frac{g^2}{2f^6 \sin^4 \theta} + \frac{f^2}{2f^4 \sin^4 \theta f^2}$$

~~$$T_{00} = \frac{1}{2f^2} \left[\frac{g^2}{x^2 f^6} + \frac{f^2}{x^4 f^2} \right]$$~~
$$= \frac{g^2}{2f^8} + \frac{f^2}{2f^6 x^4}$$

$$\begin{aligned}
 \frac{R_{tt}}{g^{tt} g} &= g_{tt} (g^{x\theta} T_{x\theta}) - 2 T_{tx} \\
 &= -\frac{x^2}{x^2 x^2} \left[\frac{f^2}{4x} + \frac{g^2}{f^4 x^4} \right] \\
 &\quad - \frac{1}{4} \left[-\frac{f^2}{2x^5} - \frac{g^2}{2f^4 x^4} \right. \\
 &\quad \left. - \frac{f^2}{2x^5 f^2} - \frac{g^2}{2x^4 f^4} \right] \\
 &\quad + \frac{g^2}{2f^{10}} + \frac{f^2}{2f^6 x^4} \\
 &\quad + \frac{g^2}{2f^8 \sin^6 \theta} + \frac{f^2}{2f^6 \sin^6 \theta x^4}
 \end{aligned}$$

$$T_{\theta\theta} = -\frac{f^2}{24} - \frac{g^2 x}{2f^4}$$

$$T_{\phi\phi} = \frac{g^2 \sin^2 \theta}{2f^2} + \frac{f^2 g^2 \sin^2 \theta}{24x}$$

$$T_{tx} = \frac{f^2}{2x} + \frac{g^2 x}{2f^4}$$

$$T_{\theta\theta} = \frac{g^2}{2f^2} + \frac{f^2 x^2}{2x^4}$$

Bianchi identities give us

$$g(f) = \text{constant}$$