

Total rate for $n \rightarrow p$ conversion per neutron

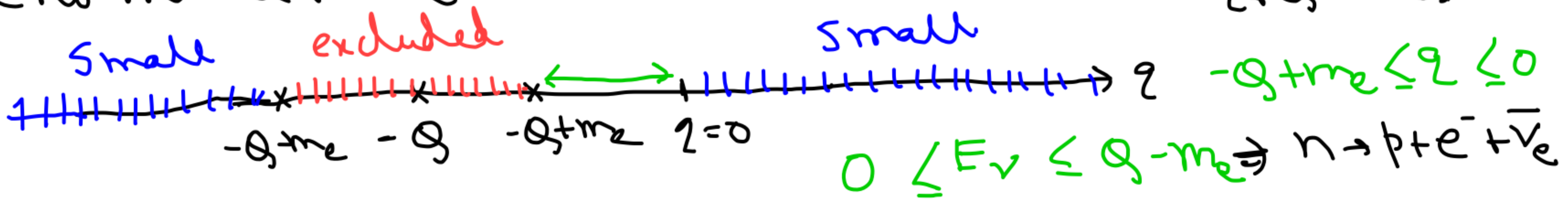
$$R(n \rightarrow p) = A \int_{-\infty}^{\infty} dq \left(1 - \frac{m_e^2}{(q+Q)^2}\right)^{1/2} \frac{q^2 (q+Q)^2}{\left(1 + e^{\frac{q}{T} - \frac{\mu_\nu}{T}}\right) \left(1 + e^{-\frac{q+Q}{T} + \frac{\mu_e}{T}}\right)}$$

$|q+Q| \geq m_e$

$R(p \rightarrow n) \rightarrow$ similar formula with the sign of the exponents reversed in the denominator.

① $T \ll Q \sim \text{MeV} \Rightarrow T_\nu \ll Q$

Denominator for $R(n \rightarrow p)$ is large. $\left. \begin{array}{l} q \gg 0 \\ q+Q \ll 0 \end{array} \right\} \text{Unit } T$



Ex. Check that for $T \ll Q \sim \text{MeV}$, $R(p \rightarrow n)$ is small (Denominator is large for all q).

② Estimate the time/energy at which the rates go below the expansion rate.

Consider $T \gtrsim \text{MeV}$.

$$\begin{aligned}
 R(n \rightarrow p) &= A \int_{-T}^T dq \left(1 - \frac{m_e^2}{(q+Q)^2}\right) \frac{q^2 (q+Q)^2}{\left(1 + e^{-\frac{q+Q}{T}}\right) \left(1 + e^{\frac{q}{T}}\right)} \\
 &\sim A \int_{-T}^T dq q^4 \sim A T^5 \\
 A &\sim (\text{Gweak})^2 \rightarrow 10^{-5} (\text{GeV})^{-2} \\
 &\rightarrow 10^{-10} (\text{GeV})^{-4} T^5 \\
 &\equiv (\text{sec})^{-1} (T/\text{MeV})^5
 \end{aligned}$$

$\frac{\dot{\chi}}{\chi} \sim T^2 \Rightarrow$

$$\frac{1}{\text{sec}} \left(\frac{T}{\text{MeV}} \right)^2$$

Recall the
neutrino decoupling
analysis

$$\Rightarrow R(n \rightarrow p) \sim \frac{1}{\text{sec}} \left(\frac{T}{\text{MeV}} \right)^5$$

falls below

the

expansion rate

when $\frac{T}{\text{MeV}} < 1$

$$\Rightarrow T < \text{MeV}.$$

Time evolution of n_n and n_p :

$$\frac{d}{dt} (n_n \lambda^3) = -R(n \rightarrow p) n_n \lambda^3 + R(p \rightarrow n) n_p \lambda^3$$

$$\frac{d}{dt} (n_n + n_p) \lambda^3 = 0$$

Define: $X_n = \frac{n_n}{n_n + n_p}$, $1 - X_n = \frac{n_p}{n_n + n_p}$

$$\Rightarrow \frac{d}{dt} X_n = -R(n \rightarrow p) X_n + R(p \rightarrow n) (1 - X_n)$$

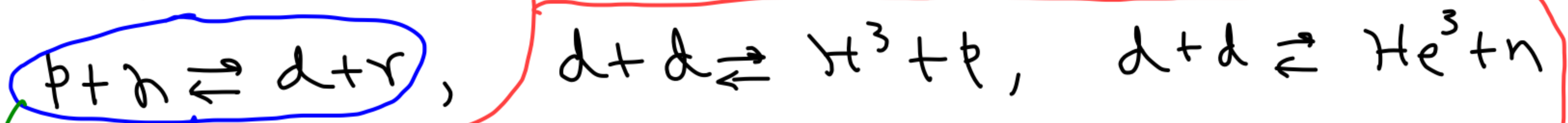
Divide by
 $(n_n + n_p) \lambda^3$

T (in K)	t (sec)	X_n
10^{11} K (8.6 MeV)	0	.4626 (stat mech)
3×10^{10} K	.101	.3798
10^{10} K	.998	.2386
3×10^9 K	12.66	.1654
10^9 K (.086 MeV)	168.1	.1333
3×10^8 K	1980	.0172
10^8 K	17780	3.09×10^{-10}

Nucleo
synthesis →

$$\frac{X_n}{1-X_n} \approx e^{-Q/T} \approx e^{-10} = .99$$

Step 3. Process of nucleosynthesis.



em interaction

not as fast as strong interaction.

Strong interaction.
fast

density of
type
nucleus.

For full analysis, define $Y_i = \frac{n_i}{n_B}$ rate of production - rate of destruction.

cross-sections are known.

We'll use a simplifying assumption that gives reasonably accurate results:

We'll estimate the error

We'll assume that nucleosynthesis begins at some time t_1 when the density of d reaches a critical value.

Once it begins it happens very fast and we can estimate the X_i 's by using equilibrium stat. mech. formula.

Two Steps.

a) Estimate the time t_i (equivalently the temperature) at which nucleosynthesis happens.

b) Use equilibrium stat mech to compute the Y_i 's.

Begin with step (b).

n_i : density of nuclei of type i

Z_i : # of protons inside i -th type nucleus

$A_i - Z_i$: # of neutrons inside i -th type "

Conserved charges:

$$(n_p)_{\text{tot}} = n_p + \sum_i Z_i n_i$$

$$(n_n)_{\text{tot}} = n_n + \sum_i (A_i - Z_i) n_i$$

$$\tilde{\mu}_1 (n_p)_{\text{tot}} + \tilde{\mu}_2 (n_n)_{\text{tot}} = \mu_p n_p + \mu_n n_n + \sum_i \mu_i n_i$$

$$\Rightarrow \mu_p = \tilde{\mu}_1, \mu_n = \tilde{\mu}_2, \mu_i = Z_i \tilde{\mu}_1 + (A_i - Z_i) \tilde{\mu}_2 = Z_i \mu_p + (A_i - Z_i) \mu_n$$

$$\mu_i = z_i \mu_p + (A_i - z_i) \mu_n, \quad y_i = n_i/n_B, \quad y_p = n_p/n_B$$

All baryons are non-relativistic. $y_n = n_n/n_B$

$$n_p = 2 e^{-\frac{m_p - \mu_p}{T}} \left(\frac{m_p T}{2\pi} \right)^{3/2} \quad T = T(t)$$

$$n_n = 2 e^{-\frac{m_n - \mu_n}{T}} \left(\frac{m_n T}{2\pi} \right)^{3/2}$$

m_n

$$n_i = g_i e^{-\frac{M_i - \mu_i}{T}} \left(\frac{M_i T}{2\pi} \right)^{3/2}$$

$A_i m_n$

$$M_i = z_i m_p + (A_i - z_i) m_n - B_i \rightarrow \text{binding energy}$$

$$\frac{y_i}{y_p^{z_i} y_n^{A_i - z_i}} = n_B^{A_i - 1} \frac{n_i}{n_p^{z_i} n_n^{A_i - z_i}} = \frac{g_i}{2^{z_i}} A_i^{3/2} e^{B_i/T} \left(\frac{T m_n}{2\pi} \right)^{A_i - 1}$$

$$Y_i = Y_p^{z_i} Y_n^{A_i - z_i} \frac{g_i}{2} A_i^{3/2} e^{B_i/T} \in A_i^{-1}$$

$\lambda T = \text{const.}$

$$\epsilon = \frac{1}{2} n_B \left(\frac{T m_p}{2\pi} \right)^{-3/2}$$

$$n_B = \frac{n_B(t_0) \lambda(t_0)^3 = 1}{\lambda(t_1)^3} = \frac{n_B(t_0)}{T_0^3} T(t_1)^3$$

$$n_B = \Omega_B \rho_c = \Omega_B \frac{3 H_0^2}{8\pi G} \quad \Omega_B \sim 0.4$$

$$H_0 = 70 \text{ km/s/L/Mpc} \rightarrow 3 \times 10^{24} \text{ cm} \Rightarrow \epsilon \sim 10^{-15}$$

$$n_p + \sum_i z_i n_i = (n_p)_{\text{tot}} = (1 - X_n(t_1)) n_B$$

$$Y_p + \sum_i z_i Y_i = 1 - X_n(t_1), \quad Y_n + \sum_i (A_i - z_i) Y_i = X_n(t_1)$$