

de Sitter metric:

$$ds^2 = -dt^2 + \lambda(t)^2 d\vec{x}^2$$

$$d\vec{x}^2 = (dx^1)^2 + (dx^2)^2 + (dx^3)^2$$

$$\lambda(t) = C e^{\mathcal{H}t}$$

Define τ such that $d\tau = \lambda(t)^{-1} dt$

$$\Rightarrow \tau = -C^{-1} \mathcal{H}^{-1} e^{-\mathcal{H}t} \quad -\infty < t < \infty$$

$$ds^2 = \mathcal{H}^{-2} \tau^{-2} (-d\tau^2 + d\vec{x}^2)$$

$$\Rightarrow -\infty < \tau < 0$$

$$\tilde{\phi} = \phi - \bar{\phi}$$

$\hookrightarrow$ average value.

$$L = \frac{1}{2} \int d^3x \mathcal{H}^{-2} \tau^{-2} \left\{ (\partial_\tau \tilde{\phi})^2 - (\vec{\nabla} \tilde{\phi})^2 \right\}$$

$$\pi(\vec{x}, \tau) = \mathcal{H}^{-2} \tau^{-2} \partial_\tau \tilde{\phi}(\vec{x}, \tau)$$

$$[\tilde{\phi}(\vec{x}, \tau), \tilde{\pi}(\vec{y}, \tau)] = i \delta^{(3)}(\vec{x} - \vec{y})$$

$$\Rightarrow [\tilde{\phi}(\vec{x}, \tau), \partial_\tau \tilde{\phi}(\vec{y}, \tau)] = i \mathcal{H}^2 \tau^2 \delta^{(3)}(\vec{x} - \vec{y})$$

$$\tau = -C^{-1} \mathcal{H}^{-1} e^{-\mathcal{H}t}$$

$$d\tau = -\mathcal{H} \tau dt$$

$$\partial_t \tilde{\phi} = -\mathcal{H} \tau \partial_\tau \tilde{\phi}$$

$$[\tilde{\phi}(\vec{x}, t), \partial_t \tilde{\phi}(\vec{y}, t)] = -i \mathcal{H}^3 \tau^3 \delta^{(3)}(\vec{x} - \vec{y})$$

$$= -i \delta^{(3)}(\mathcal{H}^{-1} \tau^{-1} (\vec{x} - \vec{y}))$$

$$\rightarrow 0 \text{ as } \tau \rightarrow \infty$$

$$[\tilde{\phi}(\vec{x}, t), \partial_t \tilde{\phi}(\vec{y}, t)] \rightarrow 0 \text{ as } t \rightarrow \infty \Rightarrow \tilde{\phi} \text{ behaves as classical field.}$$

$$\left. \begin{aligned} &\langle \tilde{\phi}(\vec{x}, t) \tilde{\phi}(\vec{y}, t) \rangle \\ &\xrightarrow{\text{large } t} \phi_a(\vec{x}, t) \phi_a(\vec{y}, t) \end{aligned} \right\} \hat{\phi}(\vec{k}, \tau) = \int d^3x e^{-i\vec{k} \cdot \vec{x}} \tilde{\phi}(\vec{x}, \tau)$$

Initial condition given by quantum result
 Recall explicit result in momentum space:

$$\langle \hat{\phi}(\vec{k}, \tau) \hat{\phi}(\vec{k}', \tau) \rangle = \frac{\mathcal{H}^2 \tau^2}{2\omega} \left(1 + \frac{1}{\omega^2 \tau^2} \right) (2\pi)^3 \delta^{(3)}(\vec{k} + \vec{k}')$$

$$\xrightarrow{\omega\tau \ll 1} \frac{\mathcal{H}^2}{2\omega^3} (2\pi)^3 \delta^{(3)}(\vec{k} + \vec{k}')$$

$\omega^{-1} \gg \tau^{-1} = \dot{\eta}^{-1} H^{-1}$ in Minkowski space: $\frac{1}{2\omega}$

Classical eq. for $\hat{\phi}$:

$$S = \frac{1}{2} \int d\tau \int \frac{d^3 k}{(2\pi)^3} \mathcal{H}^{-2} \tau^{-2} \left[\partial_\tau \hat{\phi}(\vec{k}, \tau) \partial_\tau \hat{\phi}(\vec{k}, \tau) - \vec{k}^2 \hat{\phi}(\vec{k}, \tau) \hat{\phi}(\vec{k}, \tau) \right]$$

↓ eq. of motion:

$= \omega^2$

$$\partial_\tau (\mathcal{H}^{-2} \tau^{-2} \partial_\tau \hat{\phi}) + \omega^2 \hat{\phi} \mathcal{H}^{-2} \tau^{-2} = 0$$

$$\times \mathcal{H}^2 \tau^2 \Rightarrow \partial_\tau^2 \hat{\phi} + \omega^2 \hat{\phi} - \frac{2}{\tau} \partial_\tau \hat{\phi} = 0$$

$$\Rightarrow \partial_\tau \hat{\phi} = \frac{1}{2\tau} (\partial_\tau^2 \hat{\phi} + \omega^2 \hat{\phi}) \rightarrow 0 \text{ as } \tau \rightarrow 0$$

We now repeat the analysis for metric fluctuation.

Average metric: $\overline{ds^2} = -dt^2 + \lambda(t)^2 d\vec{x}^2$

$$\lambda(t) = Ce^{Ht}.$$

Fluctuating metric:

$$ds^2 = -(1 + 2\Phi) dt^2 + 2\lambda(t) B_i dx^i dt + \lambda(t)^2 (\delta_{ij} + h_{ij}) dx^i dx^j$$

Φ, h_{ij}, B_i depend on $\vec{x}, t$.

The background metric $d\vec{s}^2$ is rotationally invariant $x^i \rightarrow R_{ij} x^j$ $RR^T = I$

$\Rightarrow$ convenient to organize the perturbations in representations of $SO(3)$.

$$B_i = \partial_i B - S_i, \quad \partial_i S_i = 0$$

Momentum space: $\hat{B}_i = ik_i \hat{B} - \hat{S}_i, \quad k_i \hat{S}_i = 0$

$$\begin{aligned} h_{ij} &= \frac{3 \times 4}{2} = 6 \\ \psi, E &: 1+1 \\ \hat{F}_i &: 2, \hat{h}_{ij}^T = 6 - 3 - 1 = 2 \end{aligned}$$

Similarly:

$$h_{ij} = -2\psi \delta_{ij} + 2\partial_i \partial_j E + (\partial_i F_j + \partial_j F_i) + h_{ij}^T$$

$$\partial_i F_i = 0 \Rightarrow k_i \hat{F}_i = 0, \quad \partial_i h_{ij}^T = 0, \quad \sum_{i=1}^3 h_{ii}^T = 0 \Rightarrow k_i \hat{h}_{ij}^T = 0$$

$$\sum_i \hat{h}_{ii} = 0.$$

Substitute into the Einstein action and expand up to quadratic order in perturbation.

Linear terms will vanish.

Quadratic terms are non-trivial.

In momentum space:

$$\frac{1}{16\pi G} \int \sqrt{-d^4x} g R$$

$$S = \int d^4x \int \frac{d^3k}{(2\pi)^3} [\dots]$$

Cannot contain $\times$ terms for different $so(3)$ representations since indices cannot contract.
 $\hat{\Phi} \hat{S}_i k_i = 0$ $h^T_{ij} \hat{S}_k \times$

We'll first analyze the action for $\hat{h}_{ij}^T$.

$\hat{h}_{ij}^T$ has 2 degrees of freedom.

Choose a basis of polarizations $e_{ij}^{(s)}, s=1,2$

$$k_i e_{ij}^{(s)}(\vec{k}) = 0, \quad \sum_{i=1}^3 e_{ii}^{(s)} = 0 \quad \text{for } s=1,2.$$

Orthonormality condition:

$$\sum_{i,j} e_{ij}^{(s)} e_{ij}^{(s')} = \delta_{ss'}$$

$$e_{ij}^{(s')} = \delta_{ss'}$$

degrees of freedom.

$$\hat{h}_{ij}^T(\vec{k}, \tau) = \sum_{s=1}^2 e_{ij}^{(s)}(\vec{k}) \hat{h}_s(\vec{k}, \tau)$$

Ex. Check that the Einstein action for tensor perturbation, reduces to ($8\pi G=1$):

$$S_{\text{tensor}} = \frac{1}{8} \int d\tau \int \frac{d^3 k}{(2\pi)^3} \mathcal{H}^{-2} \tau^{-2} \times \sum_{s=1}^2 \left[\partial_\tau \hat{h}_s(-\vec{k}, \tau) \partial_\tau \hat{h}_s(\vec{k}, \tau) - \omega^2 \hat{h}_s(-\vec{k}, \tau) \hat{h}_s(\vec{k}, \tau) \right]$$

$$\omega^2 \equiv \vec{k}^2$$

+ Cubic & higher order terms.

$\frac{1}{2} \hat{h}_s$ behaves as a scalar ϕ .

$$\frac{1}{4} \langle \hat{h}_s(\vec{k}, \tau) \hat{h}_{s'}(\vec{k}', \tau) \rangle = \delta_{ss'} \frac{(2\pi)^3}{\mathcal{H}^2 \tau^2} g^{(3)}(\vec{k} + \vec{k}') \left(1 + \frac{1}{\omega^2 \tau^2} \right)$$

Note: We have used Bunch-Davies vacuum for $\hat{h}_s$.

$$\tau \rightarrow 0 \quad \langle \hat{h}_s(\vec{k}, \tau) \hat{h}_{s'}(\vec{k}', \tau) \rangle$$
$$= (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \delta_{ss'} \frac{4H^2}{2\omega^3}$$

→ tensor perturbation.

→ produces particular type of fluctuations in CMB polarization. (Discussed later)

→ not observed yet.

For scalar and vector fluctuations we need to take into account the effect of gauge transformation.

$$d\bar{s}^2 = -dt^2 + \lambda(t)^2 d\vec{x}^2.$$

Change coordinates:

$$t \rightarrow t - \xi^0(\vec{x}, t), \quad x^i \rightarrow x^i - \xi^i(\vec{x}, t)$$

$$\xi^i = \partial_i \xi + \xi_i, \quad \partial_i \xi_i \equiv 0$$

$d\bar{s}^2 \rightarrow d\bar{s}^2 + (\dots) \Rightarrow$ Produces fake changes.

$\Rightarrow$ Some of the deformations encoded in $\Phi, \psi, E_i, F_i, S_i, B, \dots$ etc. are just gauge transformation (change of coordinates)

This does not affect h^T_{ij} since the gauge trs. parameters are scalars, or vectors.

Result: There are no dynamical scalar and vector modes in the metric.

Example: Vector perturbation:

$$t \rightarrow t, \quad x^i \rightarrow x^i - \xi^i(\vec{x}, t), \quad \partial_i \xi^i = 0$$

(Lower & upper i indices are the same)

$$d\bar{s}^2 \rightarrow -dt^2 + \lambda(t)^2 \delta_{ij} (dx^i - \partial_k \xi^i dx^k - \partial_t \xi^i dt)$$

$$(dx^j - \partial_k \xi^j dx^k - \partial_t \xi^j dt)$$

$$\stackrel{\text{ex.}}{=} d\bar{s}^2 = -2\lambda(t)^2 \partial_t \xi^i dt dx^i - \lambda(t)^2 (\partial_i \xi_j + \partial_j \xi_i) dx^i dx^j$$

Compare with:

$$ds^2 = d\bar{s}^2 + 2\lambda(t) S_i dx^i dt + \lambda(t)^2 (\partial_i F_j + \partial_j F_i) dx^i dx^j + O(\xi^2)$$

$$\Rightarrow S_i \rightarrow S_i + \lambda \partial_t \xi^i, \quad F_i \rightarrow F_i - \xi_i$$

$$S_i \rightarrow S_i + \lambda \partial_t \zeta_i, \quad F_i \rightarrow F_i - S_i$$

$\Rightarrow \left(\frac{S_i}{\lambda} + \partial_t F_i \right)$ is gauge invariant.

$\mathcal{E}_\kappa$. Check that Einstein's action for vector perturbation depends only on $\left(\frac{S_i}{\lambda} + \partial_t F_i \right)$

$\mathcal{E}_\kappa$. Linearized Einstein's eq.

$$\Rightarrow \vec{\nabla}^2 \left(\frac{S_i}{\lambda} + \partial_t F_i \right) = 0 \Rightarrow \frac{\vec{\nabla}^2}{k^2} \left(\frac{\hat{S}_i}{\lambda} + \partial_t \hat{F}_i \right) = 0$$

$$\text{For } \vec{k} \neq 0, \quad \frac{\hat{S}_i}{\lambda} + \partial_t \hat{F}_i = 0 \quad \left| \quad \int \hat{M}_i \vec{\nabla}^2 \hat{M}_i = \hat{M}_i \right.$$