

Scalar part of the metric fluctuation

$$ds^2 = -(1+2\Phi) dt^2 + 2\lambda(t) \partial_i B dx^i dt + \lambda(t)^2 (\delta_{ij} - 2\Psi \delta_{ij} + 2\partial_i \partial_j E) dx^i dx^j$$

$$t \rightarrow t - \zeta^0(\vec{x}, t), \quad x^i \rightarrow x^i - \partial_i \zeta(\vec{x}, t)$$

$$\begin{aligned} ds^2 &\rightarrow - (dt - \partial_t \zeta^0 dt - \partial_i \zeta^0 dx^i)^2 + \\ & (\lambda^2 + 2\lambda \lambda' (-\zeta^0)) (dx^i - \partial_k \partial_i \zeta dx^k - \partial_t \partial_i \zeta dt) (dx^i - \partial_k \partial_i \zeta dx^k - \partial_t \partial_i \zeta dt) \\ &= d\zeta^0{}^2 + 2 \cancel{\partial_t \zeta^0} dt^2 + 2 dt dx^i \cancel{\partial_i \zeta^0} - 2 \lambda(t)^2 \cancel{\partial_k \partial_i \zeta} dx^i dx^k \\ &\quad - 2 \lambda(t)^2 \cancel{\partial_t \partial_i \zeta} dt dx^i - 2 \lambda \lambda' \zeta^0 dx^i dx^i \\ \delta\Phi &= -\partial_t \zeta^0, \quad \delta B = \lambda^{-1} \zeta^0 = \lambda \partial_t \zeta, \quad \delta E = -\zeta^0, \quad \delta\psi = \lambda \zeta^0 \end{aligned}$$

$$\delta\Phi = -\partial_t \xi^0, \quad \delta B = \lambda^{-1} \xi^0 - \lambda \partial_t \xi, \quad \delta\psi = H \xi^0, \quad \delta E = -\xi$$

One possible gauge choice:

$$\pi = 0, \quad B = 0.$$

$$\lambda = e^{2Ht}$$

Gauge invariant combinations:

$$\Phi_B = \Phi - \frac{d}{dt} (\lambda^2 \pi - \lambda B) \quad \dot{\pi} = \frac{d}{dt} \pi$$

$$\psi_B = \psi + \lambda^2 H (\pi - \frac{B}{\lambda})$$

$$\pi - \frac{B}{\lambda} + \dot{\pi} - \frac{d}{dt} \left(\frac{B}{\lambda} - \partial_t \psi \right) = \dot{\pi} - \frac{d}{dt} \left(\frac{B}{\lambda} - \frac{\dot{\psi}}{\lambda} \right)$$

Study Linearized Einstein's eq:

Ex. in momentum space:

$$k_i \left\{ \frac{d}{dt} \hat{f}_B + \mathcal{H} \hat{\theta}_B \right\} = 0$$

$$\text{BH} \left(\frac{d \hat{f}_B}{dt} + \mathcal{H} \hat{\theta}_B \right) + \frac{1}{2} k^2 \hat{f}_B = 0$$

$$\Rightarrow \text{FH} \quad k^2 \neq 0, \quad \frac{d \hat{f}_B}{dt} + \mathcal{H} \hat{\theta}_B = 0$$

$$\Rightarrow \hat{f}_B = 0 \quad \rightarrow \quad \hat{\theta}_B = 0.$$

No dynamics in the scalar mode.

Now suppose that we also have a scalar field ϕ with constant $V(\phi)$.

Ex. In the action there will be no mixing between modes of $\hat{\phi} = \phi - \bar{\phi}$ and the metric to quadratic order.

$$\langle \hat{\phi}(\vec{k}, \tau) \hat{\phi}(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{H^2}{2\omega^3}.$$

Now consider the case of slow roll inflation.

$\phi(t) \rightarrow$ fr. of t . $\dot{\phi} \neq 0$

$$S_{\phi} = \frac{1}{2} \int d^4x \sqrt{-\det g} (-g^{\mu\nu} \partial_{\mu} \phi \partial_{\nu} \phi)$$

$$\phi = \phi + \delta\phi^2$$

$$-2g^{\mu\nu} \partial_{\mu} \phi \partial_{\nu} \delta\phi$$

$\Rightarrow \delta\phi$ begins to mix with $\psi, \pi, \pi, B.$

Linear in fluctuation

fixed by the background.

Gauge trs. law of $\vec{\phi}$:

$$\bar{\phi}(t) \rightarrow \bar{\phi}(t - \xi^0) = \bar{\phi}(t) - \xi^0 \partial_t \bar{\phi}.$$

$$\phi = \phi^1 + \phi^2 \quad \delta \phi^2 = - \xi^0 \partial_t \bar{\phi}$$

By choosing ξ^0 appropriately, we can
set ϕ^2 to 0.
 \rightarrow Scalar fluctuations can be set to 0.

In the absence of fluctuation we choose t s.t.

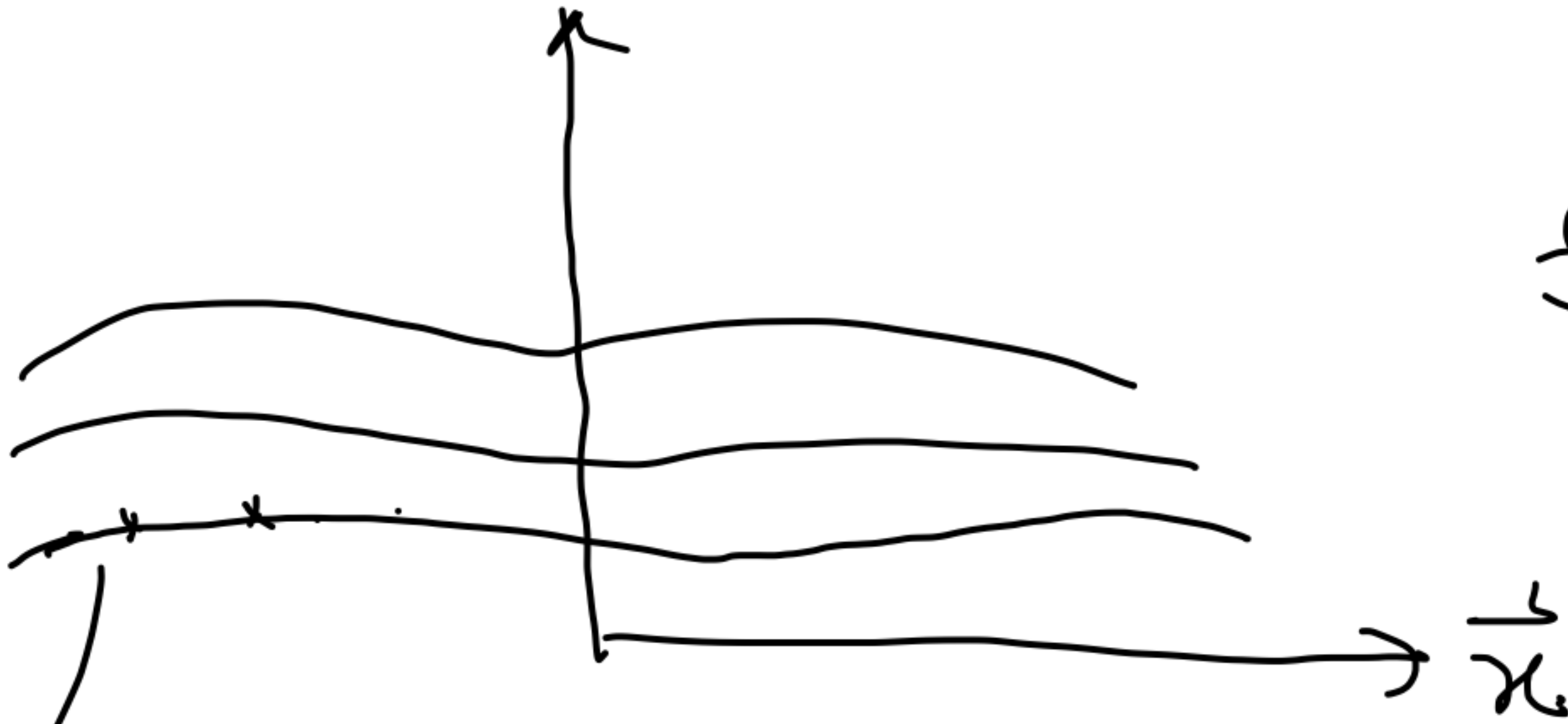
$$ds^2 = -dt^2 + \lambda(t)^2 d\vec{x}^2$$

$\phi = \text{constant}$ on a constant time slice.

$$\delta\phi = -\partial_t \phi \delta t, \quad \delta\psi = \# H \delta t$$

$\mathcal{L} = -\left(\dot{\phi} + \frac{\lambda}{\lambda} \dot{\theta}\right)$ is gauge invariant.

$$\mathcal{R} = -\mathcal{L} = \left(\dot{\phi} + \frac{\lambda}{\lambda} \dot{\theta}\right)$$



Ex. In the combined action:

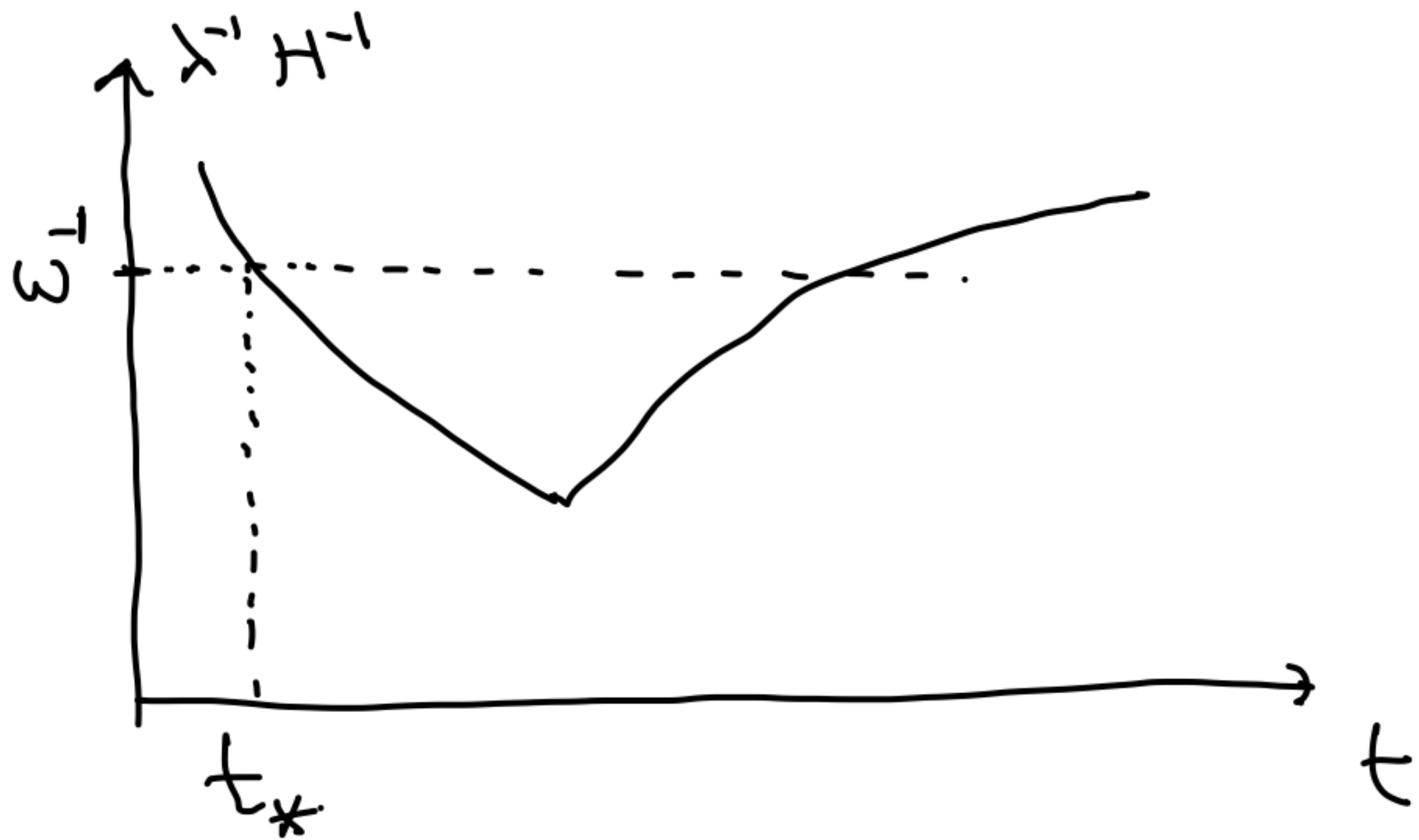
If we use $F=0$, $B=0$ gauge, then
Einstein's eq. still set $\psi=0$, $\Phi=0$

The dynamics of ϕ remains the

same as before.

$$\langle \hat{\phi}(\vec{k}, \tau) | \hat{\phi}(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{H^2}{2\omega^3}$$

$$\Rightarrow \langle \hat{R}(\vec{k}, \tau) | \hat{R}(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{H^2}{2\omega^3} \frac{H^2}{\phi^2}$$



$(1 + \frac{1}{\omega^2 \tau^2})$
 ↓
 dominates when

$$\omega \tau < 1$$

$$\omega^{-1} > \tau = \lambda^{-1} H^{-1}$$

$\dot{\phi}$ and H changes little | $H \approx H_*$
 during time H^{-1} .
 We need to calculate $\dot{\phi}$ and H at t_* .

Some definitions

$$\langle \hat{P}(\vec{k}, \tau) \hat{P}(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{\hbar^4}{2\phi^2 \omega^3}$$

P_R : scalar
power spectrum

$$\langle \hat{h}_s(\vec{k}, \tau) \hat{h}_{s'}(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{2\hbar^2}{\omega^3} \delta_{ss'}$$

$$\sum_{s=1}^2 \langle \hat{h}_s(\vec{k}, \tau) \hat{h}_s(\vec{k}', \tau) \rangle = (2\pi)^3 \delta^3(\vec{k} + \vec{k}') \frac{4\hbar^2}{\omega^3}$$

P_T : tensor
power spectrum.

Some more definitions.

$$\Delta_S^2 = \frac{\omega^3}{2\pi^2}$$

$$\mathcal{P}_R = \frac{H^4}{4\pi^2 \dot{\Phi}^2}$$

$$\Delta_T^2 = \frac{\omega^3}{2\pi^2}$$

$$\mathcal{P}_T = \frac{2H^2}{\pi^2}$$

ω -independent
 Δ_S, Δ_T is known
as scale invariant
spectrum.
→ Harrison-Zeldovich.

Δ_S, Δ_T are approximately ω independent.

t_* depends on ω via $\omega^{-1} = \lambda(t_*)^{-1} H(t_*)^{-1}$

$H, \dot{\Phi}$ depends on t_* mildly.

Deviation from scale invariance (with
dependence on ω) \rightarrow Spectral index.

to be defined later.