

~~22/11/06~~

$$G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n, \vec{g}_R, \mu) = \left\langle \prod_{i=1}^n \tilde{\Phi}_{s_i R}(k_i) \right\rangle$$

$$\left\{ \mu \frac{\partial}{\partial \mu} + \sum_{\alpha} \beta_{\alpha}(\vec{g}_R, \mu) \frac{\partial}{\partial g_{\alpha R}} + \sum_i \gamma_{s_i}(\vec{g}_R, \mu) \right\} G_{s_1 \dots s_n}^{(n)R} = 0$$

Assume : $\gamma_{s_i} = \gamma_s \delta_{s_i}$

$$G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n, \lambda g_{\alpha R}, \lambda \mu) = \lambda^{\sum d_{s_i}} \times G_{s_1 \dots s_n}^{(n)}(k_1, \dots, k_n; \{g_{\alpha R}\}, \mu)$$

where d_s : dimension of $\tilde{\Phi}_s(k)$

and d_{α} : dimension of $g_{\alpha R}$

$$g_{\alpha R} \rightarrow \lambda^{-d_{\alpha}} g_{\alpha}, \mu \rightarrow \lambda^{-1} \mu$$

$$G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n, \{g_{\alpha R}\}, \mu) = \lambda^{\sum i d_{s_i}} G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n; \{\lambda^{-d_{\alpha}} g_{\alpha}\}, \lambda^{-1} \mu)$$

We get,

$$\lambda \frac{\partial G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n, \{g_{\alpha R}\}, \mu)}{\partial \lambda}$$

$$= \left(\sum_{i=1}^n d_{s_i} - \sum_{\alpha} d_{\alpha} g_{\alpha R} \frac{\partial}{\partial g_{\alpha R}} - \mu \frac{\partial}{\partial \mu} \right) G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

~~$$\lambda^{\sum d_{s_i}} \times G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n; \{\lambda^{-d_{\alpha}} g_{\alpha}\}, \lambda^{-1} \mu)$$~~

$$G_{s_1 \dots s_n}^{(n)R}(\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

$$\lambda \frac{\partial}{\partial \lambda} G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

$$= \left(\sum_{i=1}^n d_{s_i} \left(- \sum_{\alpha} \Delta_{\alpha} g_{\alpha R} \frac{\partial}{\partial g_{\alpha R}} + \mu \frac{\partial}{\partial \mu} \right) G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu) \right) \quad \textcircled{A}$$

$$= \left[\sum_{i=1}^n \left\{ d_{s_i} + \gamma_{s_i} (\vec{f}_R, \mu) \right\} + \sum_{\alpha} \left\{ \beta_{\alpha} (\vec{f}_R, \mu) - \Delta_{\alpha} g_{\alpha R} \right\} \times \frac{\partial}{\partial g_{\alpha R}} \right]$$

$$\lambda G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu) \quad \textcircled{B}$$

eqn. where
there is no μ -derivative

- tells us how the momenta
we scale out the momenta

→ But why is this possible? → bcos
by dimensional analysis, we can transfer
the charge in momenta to other quantities

Suppose there was no need to renormalise
the theory. In this case $\beta_{\alpha} = 0, \gamma_s = 0$

The renorm. parameters/fields
can be taken to be
equal to non-renorm.
parameters/fields

$$\lambda \frac{\partial}{\partial \lambda} G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\})$$

$$= \left(\sum_{i=1}^n d_{s_i} - \sum_{\alpha} \Delta_{\alpha} g_{\alpha R} \frac{\partial}{\partial g_{\alpha R}} \right)$$

$$+ G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\})$$

(This $\overset{th}{\sim}$ should reflect the dimensional analysis)

[Suppose we change the mass scale by $(1+\epsilon)$]

Then $G_{s_1 \dots s_n}^{(n)R}$ (from the diff. case from the int. version)

$$(1+\epsilon) G_{s_1 \dots s_n}^{(n)R} \xrightarrow{\text{contract R}} (1+\epsilon) \partial k_1, \dots, (1+\epsilon) \partial k_n, (1+\epsilon) \frac{\partial}{\partial R}$$



$$= (1+\epsilon) \sum ds_i G_{s_1 \dots s_n}^{(n)R} (k_1, \dots, k_n, g_{RR})$$

Expand out both sides upto $O(\epsilon)$

$$\Rightarrow \left(\epsilon \lambda \frac{\partial}{\partial \lambda} + \sum \Delta_\alpha \epsilon g_{RR} \frac{\partial}{\partial f_{\alpha R}} \right) G = \epsilon \sum ds_i G$$

$$\Rightarrow \frac{\partial}{\partial \lambda} G = \left(\sum ds_i - \sum \Delta_\alpha g_{\alpha} \frac{\partial}{\partial f_\alpha} \right) G$$

This relation is the same as the special case of (B), when we don't need to renormalize.

from (C) of previous copy, we get RATE :-

$$(1+\epsilon) \sum ds_i \xrightarrow{\text{MIR}} G_{s_1 \dots s_n}^{(n)R} (k_1, \dots, k_n, g_{RR}) \lambda^{-4\alpha} (1-\beta\mu)^{\alpha} \text{ (inter)} \\ \text{(though (A) is qualitatively a diff. eqn than (B))}$$

the one we obt. here, by renom. grp. eqns, we converted the μ -deriv. term into other terms

$$= (1+\epsilon) \sum ds_i \left[G_{s_1 \dots s_n}^{(n)R} (k_1, \dots, k_n, g_{RR}), \lambda^{-4\alpha}, \mu^{-1} \right] + O(\epsilon) \\ = (1+\epsilon) \sum ds_i G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n, \{g_{\alpha\beta}\}, \mu) + O(\epsilon)$$

(B) \Rightarrow generalization of the case when we don't need to renormalize the theory

$ds_i + \gamma_i$ doesn't mean that the field has charged dimensions, it is as if we need to assign a dim. $ds_i + \gamma_i$ to the fields if we want to arrive at this eqn. by dimensional analysis — this additional term is coming from μ -deriv. elimination.

In the more general case we get an anomalous dim. matrix which we will have to diagonalise going to an appropriate basis.

$$ds_i + \gamma_{s_i} \rightarrow$$

↓

as if $\phi_{\lambda R}$ has dimension

$$ds + \gamma_s(\vec{g}_R, \mu)$$

↓

anomalous dimension

so we have $\lambda \frac{\partial}{\partial \lambda} G^{(n)R}_{s_1 \dots s_n}(\lambda k_1, \dots, \lambda k_n, \vec{g}_R, \mu)$

$$= \left[\sum_{i=1}^m ds_i + \gamma_{s_i}(\vec{g}_R, \mu) \right] + \left[\sum_{\alpha} [\beta_{\alpha}(\vec{g}_R, \mu) - \lambda g_{\alpha R}] \frac{\partial}{\partial g_{\alpha R}} \right]$$

$\times G^{(n)R}_{s_1 \dots s_n}(\lambda k_1, \dots, \lambda k_n, \vec{g}_R, \mu) \quad \boxed{\beta_{\alpha}(\vec{g}_R, \mu)}$

We would want to int. this diff. eqn w.r.t. λ & express $G^{(n)R}(k_i)$ to

$$G^{(n)R}(k_i) \rightarrow \lambda = 1$$

Solution :- Define : $\vec{g}_{\alpha R}(\lambda, \vec{g}_R)$ to be the solution to

$$\lambda \frac{\partial}{\partial \lambda} \vec{g}_{\alpha R}(\lambda, \vec{g}_R) = \tilde{\beta}_{\alpha}(\vec{g}_{\alpha R}(\lambda, \vec{g}_R))$$

$$\vec{g}_{\alpha R}(\lambda, \vec{g}_R) = g_{\alpha R} \text{ at } \lambda = 1$$

where
 $\tilde{\beta}_{\alpha} = \beta_{\alpha} - \lambda g_{\alpha R}$

(Dep. of $\bar{g}_{\lambda R}$ on λ comes bcos it satisfies
a diff. eqn. in $\lambda \Rightarrow$

$$\lambda \frac{\partial}{\partial \lambda} \bar{g}_{\lambda R} = P_{\lambda}(\bar{g}_R) \rightarrow \text{so } \bar{g}_{\lambda R} \text{ depends on } \lambda$$

($\bar{g}_{\lambda R}$ depends on \bar{g}_R through boundary condns; otherwise the above diff. eqn. knows nothing about $\bar{g}_R \Rightarrow$)

$$\bar{g}_{\lambda R} \text{ at } \lambda=1 = g_{\lambda R} \rightarrow \text{Boundary condn.}$$

Example \Rightarrow suppose $P(\bar{g}_R, \mu) = -A \bar{g}_R^3$

$$\therefore \lambda \frac{\partial}{\partial \lambda} \bar{g}_R = -A \bar{g}_R^3$$

$$\Rightarrow (\text{It is the same as}) \lambda \frac{\partial}{\partial \lambda} \left(\frac{1}{2 \bar{g}_R^2} \right) = A$$

$$\Rightarrow \frac{1}{2 \bar{g}_R^2} = A \ln \lambda + C$$

↓
integration constant

(Apply B. c. \Rightarrow)

$$\bar{g}_R = g_R \text{ at } \lambda = 1$$

$$\Rightarrow C = \frac{1}{2 g_R^2}$$

$$\therefore \frac{1}{2 \bar{g}_R^2} = A \ln \lambda + \frac{1}{2 g_R^2} = \frac{1 + 2 A g_R^2 \ln \lambda}{2 g_R^2}$$

$$\boxed{\frac{\bar{g}_R^2}{g_R^2} = \frac{g_R^2}{1 + 2 A g_R^2 \ln \lambda}}$$

$$\bar{g}_R(\lambda, g_R)$$

In other
cases, suppose
 μ -dep'ndence
is no μ -deriv.,
that we take μ
to be fixed

[We need one initial card per variable
bcos we have first order diff. eqns — this
is supplied by \vec{g}_R]

CLAIM \Rightarrow

Solution for $Q_{s_1, \dots, s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n, \vec{g}_R, u)$

$$= Q_{s_1, \dots, s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \vec{g}_{\text{far}} (\lambda, \vec{g}_R))$$

$$\exp \left[\sum_{i=1}^n ds_i \ln \lambda \right]$$

$$\times \exp \left[\sum_{i=1}^n \int_1^\lambda \frac{d\lambda'}{\lambda'} \gamma_{s_i} (\{ g_\alpha (\lambda', \vec{g}_R) \}) \right] \quad \textcircled{C}$$

↓
Drop this out
i.e., & take this
equal to some
fixed const.

(We'll do the proof later but we can do some)

Consistency check \rightarrow

Take the case when we don't need to renormalise the theory

Suppose $\gamma_{s_i} = 0, \beta_\alpha = 0$

Then,

$$\tilde{\beta}_\alpha = -\Delta_\alpha \vec{g}_{\text{far}}; \lambda \frac{\partial}{\partial \lambda} \vec{g}_{\text{far}} = -\Delta_\alpha \vec{g}_{\text{far}}$$

$$[\text{Bcos } \tilde{\beta}_\alpha = \beta_\alpha - \Delta_\alpha \vec{g}_{\text{far}}] \Rightarrow \boxed{\vec{g}_{\text{far}} = \lambda^{-\Delta_\alpha} \vec{g}_{\text{far}}}$$

From (C) we have $G_{s_1, \dots, s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \vec{g}_R)$

$$= Q_{s_1, \dots, s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \lambda^{-\Delta_\alpha} \vec{g}_{\text{far}}) \lambda^{\sum ds_i}$$

\rightarrow this is precisely the result from dimensional analysis \rightarrow so consistent

[Once we know the β 's, we know γ 's & \overline{g}_α 's — so we know $\overline{g}^{(n)R}$ at the scaled momenta with coupling const. \overline{f}_R in terms of $G^{(n)R}$ at $\lambda = 1$ with coupling const.

$$[\overline{g}_{\alpha R}]$$

As $\lambda \rightarrow \infty$, on the RHS we have amplitudes at finite momentum but at coupling constant $\overline{g}_{\alpha R}(\lambda, \overline{g}_R)$ for large λ .

~~for~~ Non-abelian gauge theories:-

$$\overline{g}_R(\lambda, g_R) \rightarrow 0 \text{ for } \lambda \rightarrow \infty \text{ if } A > 0$$

$$\overline{g}_R^2 = \frac{g_R^2}{1 + 2A g_R^2 \ln \lambda}$$

Small mom. means large distance scale
→ pert. th. can't be applied

e.g. → What is the force if we want to pull apart 2 fermions (or constituents)?

→ These questions can't be addressed b/c coupling const. becomes large.
—these ques. are imp. for the formation of bound states

Perturbation theory becomes more accurate

otherwise we can't apply pert. theory

ASYMPTOTIC FREEDOM

Running of \overline{g}_R is not important for calculation of gauge invariant quantities.

But running of \overline{m}_R is important.

$$\Rightarrow \frac{d\overline{m}_R}{d\lambda} \underset{\lambda \rightarrow \infty}{\sim} f_m(\overline{m}_R, \overline{g}_R)$$

$$\Rightarrow \lambda \frac{\partial \bar{m}_R}{\partial \lambda} = (\beta_m(\bar{m}_R, \bar{g}_R) - \bar{m}_R)$$

\Downarrow

$$\sim \frac{1}{\bar{g}_R^2} \bar{m}_R$$

(can be ignored compared to \bar{m}_R for small \bar{g}_R)

$$\Rightarrow \lambda \frac{\partial \bar{m}_R}{\partial \lambda} \simeq -\bar{m}_R$$

Now, $\bar{m}_R \simeq \frac{m_R}{\lambda}$ since at $\lambda=1$ $\bar{m}_R = m_R$.

\therefore as $\lambda \rightarrow \infty$, $\bar{m}_R \rightarrow 0$

\rightarrow This can cause potential IR divergence on the right-hand side.

\rightarrow Then the RG eqns are not very useful to get large λ behaviour.

$\#$ We need to look for quantities which remain finite as $\bar{m}_R \rightarrow 0$

(Infrared safe quantities)

e.g. \rightarrow Deep inelastic scattering amplitude

Sum over all possible/large no. of possible final states
But taking a single proton, etc. will have IR div.

effective mass of fermion going to zero

this can cause problem bcs div. may come from $\bar{p}=0$ region
 \rightarrow infrared divergence

RHS = zero times infinity is not of much use (over $\bar{g}_R = 0$)

for quantities

as $\bar{m}_R \rightarrow 0$

\rightarrow these corr. do free field theory

~~23/11/06~~

$$\lambda \frac{\partial}{\partial \mu} G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

$$= \left\{ \sum_{i=1}^n (ds_i + \gamma_{s_i}(\vec{g}, \mu)) + \sum_{\alpha} \tilde{\beta}_{\alpha}(\vec{g}_R, \mu) \frac{\partial}{\partial g_{\alpha R}} \right\}$$

$$G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

where $\tilde{\beta}_{\alpha} = \beta_{\alpha} - \delta_{\alpha} g_{\alpha R}$ (no sum over α)

Solution :-

$$G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{g_{\alpha R}\}, \mu)$$

$$= \exp \left(\sum_{i=1}^n ds_i \ln \lambda \right) \exp \left(\sum_{i=1}^n \int \frac{d\lambda'}{\lambda'} \gamma_{s_i} \left\{ \bar{g}_{\alpha R}(\lambda, \vec{g}_R, \mu), \mu \right\} \right)$$

$$G_{s_1 \dots s_n}^{(n)R} (\lambda k_1, \dots, \lambda k_n; \{ \bar{g}_{\alpha R}(\lambda, \vec{g}_R, \mu) \}, \mu)$$

where $\bar{g}_{\alpha R}(\lambda, \vec{g}_R, \mu)$ is defined through:

$$\lambda \frac{\partial}{\partial \lambda} \bar{g}_{\alpha R}(\lambda, \vec{g}_R, \mu) = \tilde{\beta}_{\alpha}(\bar{g}_{\alpha R}(\lambda, \vec{g}_R, \mu), \mu)$$

$$\& \bar{g}_{\alpha R}(1, \vec{g}_R, \mu) = g_{\alpha R} \quad (\text{B.C.})$$

simplify notation :-

- ① Drop all μ factors. (but this eqn. doesn't involve any deriv. of μ ; but all fns appearing here can, in principle, depend on μ)

$$\textcircled{2} \quad \underbrace{G}_{\lambda_1, \dots, \lambda_n}^{(n)} \rightarrow G$$

$$\textcircled{3} \quad \partial_\lambda G(\lambda \kappa_1, \dots, \lambda \kappa_n; \vec{g}_R) = \frac{\partial}{\partial g_{\alpha R}} G(\kappa_1, \dots, \kappa_n; \vec{g}_R)$$

↳ derivative w.r.t. coupling const. argument

*We will
use
this
identity*

$$\text{Identity: } \sum \frac{\partial \bar{g}_{\alpha R}}{\partial g_{\alpha R}} \tilde{P}_\alpha(\vec{g}_R) = \tilde{P}_\alpha(\vec{g}_R)$$

Proof:

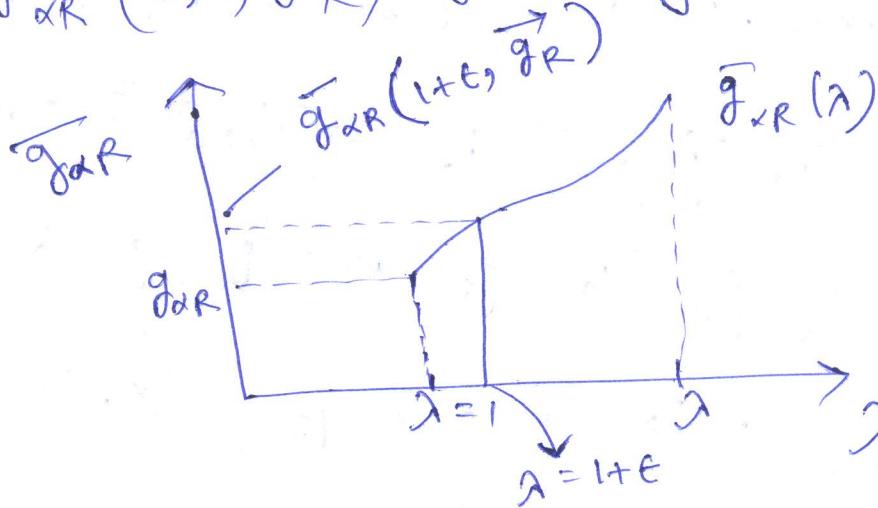
Begin with the equation

$$\lambda \frac{\partial}{\partial \lambda} \bar{g}_{\alpha R}(\lambda, \vec{g}_R) = \tilde{P}_\alpha(\vec{g}_\alpha(\lambda, \vec{g}_R))$$

(this eqn is scale invariant in λ) \rightarrow diff. eqn. is unchanged under a rescaling of λ .

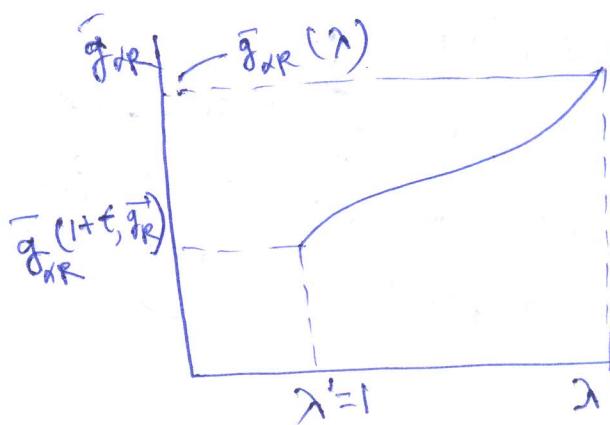
If $\bar{g}_{\alpha R}(\lambda, \vec{g}_R)$ is a poly., so is

$\bar{g}_{\alpha R}(c\lambda, \vec{g}_R)$ for any constant c .



Also, can be thought of as evolving from $\lambda = 1 + \epsilon$ & $\bar{g}_{\alpha R}(1+\epsilon, \vec{g}_R)$

$$\text{Define: } \lambda' = \frac{\lambda}{1+\epsilon}$$



Here the y-axis hasn't changed but the x-axis has changed

$$\lambda' = \frac{\lambda}{1+\epsilon}$$

(Forget about the first graph & consider the 2nd graph only - as if we are solving the diff. eqn. in the x' -plane) \rightarrow

We get the identity (evolving from 2

$$\bar{g}_{\alpha R}(\lambda) = \bar{g}_{\alpha R}\left(\frac{\lambda}{1+\epsilon}, \bar{g}_{\alpha R}(1+\epsilon, \bar{g}_R)\right)$$

diff. initial pts
~~lead~~ to the same pt.)

$$\simeq \bar{g}_{\alpha R}(\lambda - \epsilon, \bar{g}_{\alpha R}(1, \bar{g}'_R))$$

$$+ \epsilon \left. \frac{\partial}{\partial \lambda} \bar{g}_{\alpha R}(\lambda, \bar{g}_R) \right|_{\lambda=1}$$

(Taylor series exp. the argument)

$$\Rightarrow \bar{g}_{\alpha R}(\lambda) = \bar{g}_{\alpha R}(\lambda - \epsilon, \bar{g}_{\alpha R} + \epsilon \tilde{P}_x(\bar{g}'_R))$$

~~$\Rightarrow \bar{g}_{\alpha R}(\lambda, \bar{g}_R) = \bar{g}_{\alpha R}$~~

$$\Rightarrow \bar{g}_{\alpha R}(\lambda, \bar{g}_R)$$

$$= \bar{g}_{\alpha R}(\lambda, \bar{g}_R) - \epsilon \lambda \frac{\partial}{\partial \lambda} \bar{g}_R(\lambda, \bar{g}_R)$$

$$+ \epsilon \tilde{P}_x(\bar{g}_R) \frac{\partial \bar{g}_{\alpha R}(\lambda, \bar{g}_R)}{\partial \bar{g}_{\alpha R}}$$

$$= \bar{g}_{\alpha R}(\lambda, \bar{g}_R) - \epsilon \tilde{P}_x(\bar{g}_R(\lambda, \bar{g}_R))$$

$$+ \epsilon \tilde{P}_x(\bar{g}_R) \frac{\partial \bar{g}_{\alpha R}(\lambda, \bar{g}_R)}{\partial \bar{g}_{\alpha R}}$$

(Compare L.H.S. & R.H.S.) \rightarrow

Coeffs. of $O(\epsilon)$ terms give the reqd. identity.

Using
 $\bar{g}_{\alpha R}(1, \bar{g}_R, \epsilon) = \bar{g}_{\alpha R}$
 \downarrow
 $\text{But } \tilde{P}_x(\bar{g}_R)|_{\lambda=1} = \tilde{P}_x(\bar{g}_R)$
 $\&$
 $\lambda \frac{\partial}{\partial \lambda} \bar{g}_{\alpha R} = \tilde{P}_x(\lambda)$

(Proved)

LHS is encoding the charge in initial condns while the RHS involves $\lambda \frac{\partial}{\partial \lambda} \vec{g}_{\alpha R}$

Why these 2 be related?

→ We can compensate for the charge in initial condns

evolve over a diff. range

of λ — since it is scale invariant

in λ → on one end you are changing

λ & on the other, you are "the"

initial condns → these 2 are correlated

(e.g. → Evolve for 58 sec. instead of 1 min.
→ evolve from 2 sec. initial condns)

$$\sum_{\gamma} \frac{\partial \vec{g}_{\alpha R}(\lambda, \vec{g}_R)}{\partial \vec{g}_{\alpha R}} \tilde{P}_{\gamma}(\vec{g}_R) = \tilde{P}_{\alpha}(\vec{g}_R(\lambda, \vec{g}_R))$$

For the solution given, we calculate the l.h.s. of the diff. eqn which is

$$\lambda \frac{\partial}{\partial \lambda} G(x_k, \dots, x_n, \vec{g}_R))$$

$$= \left\{ \left(\sum_i dx_i \right) + \sum_i \gamma_{si} (\vec{g}_R(\lambda, \vec{g}_R)) \right\}_{G^{(1)}}$$

Analogy → consider an eqn which is time-translation invariant

~~$$\text{L.H.S.} \quad \frac{\partial}{\partial \lambda} G(\lambda k_1, \dots, \lambda k_n; \vec{g}_R)$$~~

$$= \exp\left(\sum_{i=1}^n d_{S_i} \ln \lambda\right) \exp\left(\sum_{i=1}^n \int_1^{\lambda} \frac{d\lambda'}{\lambda'} \gamma_{S_i}\left(\vec{g}_{\alpha R}(\lambda', \vec{g}_R)\right)\right)$$

$$\times \left[\left(\sum_{i=1}^n d_{S_i} + \sum_{i=1}^n \gamma_{S_i}(\vec{g}_R(\lambda, \vec{g}_R)) \right) \right]$$

$$\times G(k_1, \dots, k_n; \vec{g}_R(\lambda, \vec{g}_R))$$

$$+ \partial_\lambda G(k_1, \dots, k_n; \vec{g}_R(\lambda, \vec{g}_R)) \times \lambda \frac{\partial \vec{g}_{\alpha R}(\lambda, \vec{g}_R)}{\partial \lambda}$$

replace by
 $\tilde{\beta}_R(\vec{g}_R(\lambda, \vec{g}_R))$

R.H.S. of the differential eqn:

$$= \exp\left(\sum_i d_{S_i} \ln \lambda\right) \exp\left(\sum_{i=1}^n \int_1^\lambda \frac{d\lambda'}{\lambda'} \gamma_{S_i}\left(\vec{g}_{\alpha R}(\lambda', \vec{g}_R)\right)\right)$$

$$\times \left[\sum_{i=1}^n \left(d_{S_i} + \gamma_{S_i}(\vec{g}_R, \mu) \right) G(k_1, \dots, k_n; \vec{g}_R(\lambda, \vec{g}_R)) \right]$$
~~$$+ G(k_1, \dots, k_n; \vec{g}_R) \sum_{i=1}^n \int_1^\lambda \frac{d\lambda'}{\lambda'} \partial_\lambda \gamma_{S_i}\left(\vec{g}_{\alpha R}(\lambda', \vec{g}_R)\right) \frac{\partial \vec{g}_{\alpha R}(\lambda', \vec{g}_R)}{\partial g_{\alpha R}}$$~~

$$+ \partial_\lambda G(k_1, \dots, k_n; \vec{g}_R(\lambda, \vec{g}_R)) \frac{\partial \vec{g}_{\alpha R}(\lambda, \vec{g}_R)}{\partial g_{\alpha R}} \tilde{\beta}_R(\vec{g}_R)$$

(deriv. from the
2nd exponential)

(Using the identity we proved, we have)

$$\begin{aligned}
 & \text{RHS} \\
 &= \exp\left(\sum_i d_{s_i} \ln \lambda\right) \exp\left(\sum_i \int_1^{\lambda} \frac{dx'}{\lambda'} \gamma_{s_i}\left(\vec{g}_R(x'), \vec{g}_R\right)\right) \\
 &\quad \left[Q(k_1, \dots, k_n; \vec{g}_R(x, \vec{g}_R)) \right. \\
 &\quad \times \sum_{i=1}^n \left(d_{s_i} + \gamma_{s_i}(\vec{g}_R) \right) + \int_1^{\lambda} \frac{dx'}{\lambda'} \partial_{x'} \gamma_{s_i}(\vec{g}_R(x', \vec{g}_R)) \right. \\
 &\quad \left. \times \tilde{P}_{\gamma}(\vec{g}_R(x')) \right. \\
 &\quad \left. + \partial_{\gamma} Q(k_1, \dots, k_n; \vec{g}_R(x, \vec{g}_R)) \times \tilde{P}_{\gamma}(\vec{g}_R(x, \vec{g}_R)) \right]
 \end{aligned}$$

Everything on RHS & LHS matches
except for

$$\begin{aligned}
 & \sum_i \gamma_{s_i}(\vec{g}_R(x, \vec{g}_R)) \text{ on LHS } + \\
 & \sum_i \gamma_{s_i}(\vec{g}) + \sum_i \int_1^{\lambda} \frac{dx'}{\lambda'} \partial_{x'} \gamma_{s_i}(\vec{g}_R(x', \vec{g}_R)) \tilde{P}_{\gamma}(\vec{g}_R(x')) \\
 & \text{on the RHS}
 \end{aligned}$$

$$\text{Now, } \tilde{P}_{\gamma}(\vec{g}_R(x')) = \lambda' \frac{\partial}{\partial x'} \vec{g}_R(x', \vec{g}_R)$$

$$\begin{aligned}
 & \therefore \int_1^{\lambda} \frac{dx'}{\lambda'} \partial_{x'} \gamma_{s_i}(\vec{g}_R(x', \vec{g}_R)) \tilde{P}_{\gamma}(\vec{g}_R(x')) \\
 &= \int_1^{\lambda} dx' \frac{\partial}{\partial x'} \gamma_{s_i}(\vec{g}_R(x', \vec{g}_R)) \\
 &= \gamma_{s_i}(\vec{g}_R(x, \vec{g}_R)) - \gamma_{s_i}(\vec{g}_R(1, \vec{g}_R))
 \end{aligned}$$

(Using the identity we proved, we have)

$$\begin{aligned}
 & \underset{\text{RHS}}{=} \exp \left(\sum_i d_{S_i} \ln \lambda \right) \exp \left(\sum_i \int_1^\lambda \frac{d\lambda'}{\lambda} \gamma_{S_i} \left(\vec{g}_R(\lambda'), \vec{g}_R \right) \right) \\
 & \quad \left[Q(k_1, \dots, k_n; \vec{g}_R(x, \vec{g}_R)) \right. \\
 & \quad \times \sum_{i=1}^n \left(d_{S_i} + \gamma_{S_i} \left(\vec{g}_R \right) \right) + \int_1^\lambda \frac{dx'}{\lambda} \partial_{x'} \gamma_{S_i} \left(\vec{g}_R(x', \vec{g}_R) \right) \\
 & \quad \left. \times \tilde{P}_g \left(\vec{g}_R(x') \right) \right] \\
 & \quad + \partial_g Q(k_1, \dots, k_n; \vec{g}_R(x, \vec{g}_R)) \times \tilde{P}_g \left(\vec{g}_R(x, \vec{g}_R) \right)
 \end{aligned}$$

Everything on RHS & LHS matches
except for

$$\begin{aligned}
 & \sum_i \gamma_{S_i} \left(\vec{g}_R(x, \vec{g}_R) \right) \text{ on LHS &} \\
 & \sum_i \gamma_{S_i} \left(\vec{g} \right) + \sum_i \int_1^\lambda \frac{dx'}{\lambda} \partial_{x'} \gamma_{S_i} \left(\vec{g}_R(x', \vec{g}_R) \right) \tilde{P}_g \left(\vec{g}_R(x') \right) \\
 & \text{on the RHS}
 \end{aligned}$$

$$\text{Now, } \tilde{P}_g \left(\vec{g}_R(x') \right) = \lambda' \frac{\partial}{\partial x'} \vec{g}_R(x', \vec{g}_R)$$

$$\begin{aligned}
 & \therefore \int_1^\lambda \frac{dx'}{\lambda} \partial_{x'} \gamma_{S_i} \left(\vec{g}(x, \vec{g}_R) \right) \tilde{P}_g \left(\vec{g}_R(x') \right) \\
 & = \int_1^\lambda \frac{dx'}{\lambda} \frac{\partial}{\partial x'} \gamma_{S_i} \left(\vec{g}_R(x', \vec{g}_R) \right) \\
 & = \gamma_{S_i} \left(\vec{g}_R(x, \vec{g}_R) \right) - \gamma_{S_i} \left(\vec{g}_R(1, \vec{g}_R) \right)
 \end{aligned}$$

$$= \gamma_{S_1}(\vec{g}_R(\lambda, \vec{g}_R)) - \gamma_{S_1}(\vec{g}_R)$$

Now

this \downarrow cancels $\sum_i \gamma_{S_1}(\vec{g}_R)$

So

$$\boxed{\text{LHS} = \text{RHS}}$$

(proved)

24 | 11 | 06

Spontaneous symmetry breaking

$$S = \int d^4x \left[-\frac{1}{2} \eta^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V(\phi) \right]$$

Suppose $V(\phi) = \frac{1}{2} m^2 \phi^2 + \frac{\lambda}{4!} \phi^4$

we have taken
 λ_4 instead
of $\frac{1}{4!}$ for
convenience

Here, $S[\phi] = S[-\phi]$

$$\Rightarrow \langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$= \frac{S[\delta\phi] e^{iS[\phi]} \phi(x_1) \dots \phi(x_n)}{S[\delta\phi] e^{iS[\phi]}}$$

(change of variable) $\rightarrow \phi(x) = -X(x)$: change of variable

$$[\delta\phi] = [\delta X] \times \text{sign}$$

$$S[\phi] = S[X] \text{ due to symmetry}$$

$$\langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$= \frac{S[\delta X] e^{iS[X]} (-1)^n X(x_1) \dots X(x_n)}{S[\delta X] e^{iS[X]}}$$

↓ rename X as ϕ

$$= (-1)^n \frac{S[\delta\phi] e^{iS[\phi]} \phi(x_1) \dots \phi(x_n)}{S[\delta\phi] e^{iS[\phi]}}$$

$$= (-1)^n \langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$\Rightarrow \langle \phi(x_1) \dots \phi(x_n) \rangle = (-1)^n \langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$\Rightarrow \langle \phi(x_1) \dots \phi(x_n) \rangle = 0 \text{ if } n \text{ is odd}$$

Take a new action

$$S = \int d^4x [-\frac{1}{2} \eta^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - V[\phi]]$$

$$\text{where } V[\phi] = -\frac{1}{2} \mu^2 \phi^2 + \lambda \frac{1}{4} \phi^4$$

(We have replaced) $m^2 \xrightarrow{\text{by}} -\mu^2$

Naive pert. of mass
leads to
inconsistencies

Perturbation theory in \mathcal{D} :

$\lambda = 0 \Rightarrow$ free field theory

Single particle state $|\vec{p}\rangle$

has $H|\vec{p}\rangle = \sqrt{\vec{p}^2 - \mu^2} |\vec{p}\rangle$

→ replaces

$$\sqrt{\vec{p}^2 + m^2} |\vec{p}\rangle$$

$\sqrt{\vec{p}^2 - \mu^2}$ tells us that as if we have a particle of -ve (mass)².

Velocity $v = \frac{dp}{dt} = \frac{|\vec{p}|}{\sqrt{\vec{p}^2 - \mu^2}} > 1$

⇒ particle propagates faster than the speed of light → **TACHYONS**
→ not a consistent theory

this kind of particles are called tachyons

Physical origin of the problem

Perturbation theory is valid if ϕ is small

so that $|\frac{\lambda \phi^4}{4}| \ll |\mu^2 \phi^2|$

(To see if the above is a consistent assumption)

Consider a small fluctuation in ϕ

$$(B.C. \rightarrow \phi(\vec{x}, t) = \epsilon \text{ at } t=0, \dot{\phi}(\vec{x}, t) = 0 \text{ at } t=0)$$

(Such fluctuations of ϕ will be there bcos of quantum fluctuations ~~are~~ typically such fluctuations are small - we will see the evolution acc. to classical eqns.)

Classical eqn. of motion of ϕ :

$$(\partial_t^2 - \vec{\nabla}^2 - \mu^2) \phi = 0 \text{ for small } \gamma$$

similar to K.G. eqn except for $-\mu^2$ in place of $+m^2$

For space-independent ϕ , $(\partial_t^2 - \mu^2)\phi = 0$

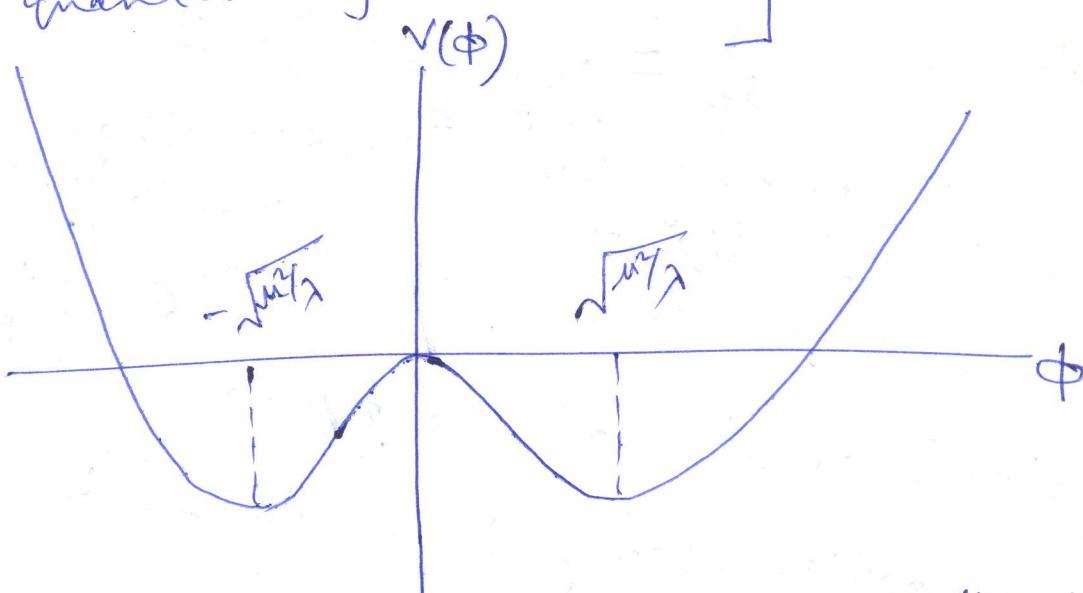
$$\phi = A e^{i\omega t} + B e^{-i\omega t}$$

11 b.c.

$$\epsilon \cosh i\omega t \rightarrow \infty \text{ as } t \rightarrow \infty$$

Fair $\phi = A e^{i\omega t} + B e^{-i\omega t} \stackrel{b.c.}{=} \epsilon \cosh i\omega t \rightarrow \infty \text{ as } t \rightarrow \infty$

[∴ taking ϕ to be small is not a consistent assumption - so until you take into account the effect of $\propto \phi^4$, ϕ grows uncontrollably - ∴ ϕ^4 term becomes as imp. as ϕ^2 term bcos of quantum fluctuations]

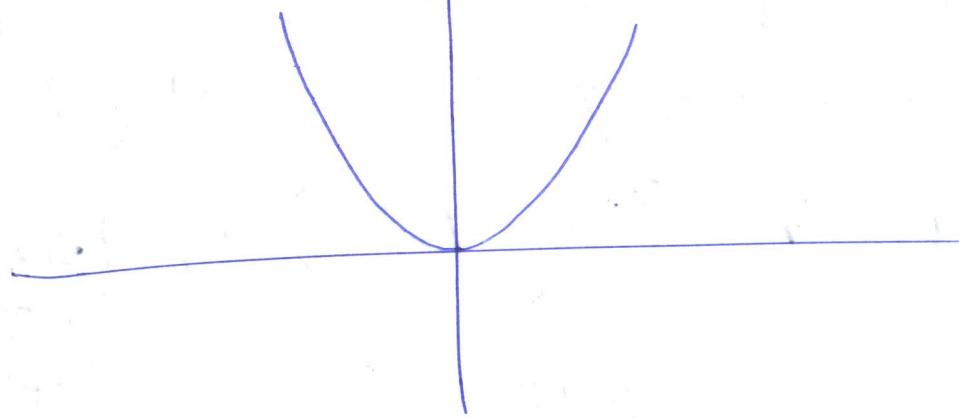


(Motion of ϕ will be the same as a particle in this potential)

(pert. th. isn't good here → consider a slight displacement from the origin & ...)

Contrast
this case
with

$$\frac{1}{2}m\phi^2 + \lambda/4\phi^4 \quad m, \lambda > 0$$

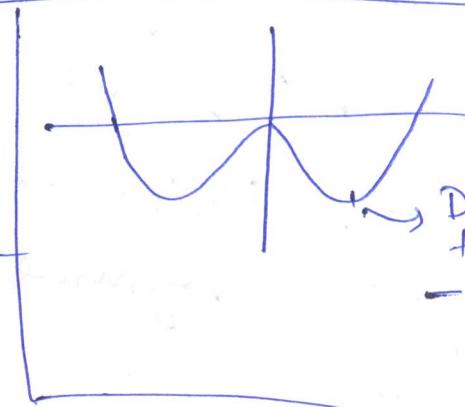


(pert. th.
is good
here, bcs
displacement
from origin
---)

Remedy :-

Carry out perturbation theory around a minimum of $V(\phi)$.

$$V(\phi) = -\frac{1}{2}\mu^2\phi^2 + \lambda/4\phi^4$$



Do pert.
th. here
- small
fluctuation
→ consistent

$$V'(\phi) = -\mu^2\phi^2 + \lambda\phi^3$$

$$\therefore V'(\phi) = 0 \Rightarrow \phi^2 = \mu^2/\lambda, \phi = \pm\sqrt{\mu^2/\lambda}$$

[we can do pert. at $\phi = \sqrt{\mu^2/\lambda}$ or at $\phi = -\sqrt{\mu^2/\lambda}$ \Rightarrow 2 choices
→ let's choose $\phi = \sqrt{\mu^2/\lambda}$]

Define a new field $X = \phi - \sqrt{\mu^2/\lambda}$

$$\therefore \phi = X + \sqrt{\mu^2/\lambda}$$

$$\begin{aligned} \therefore V(\phi) &= -\lambda/4\left(\frac{\mu^2}{\lambda}\right)^2 + \mu^2X^2 + \lambda\sqrt{\mu^2/\lambda}X^3 \\ &\quad + \lambda/4X^4 \\ &\equiv \tilde{V}(X) \end{aligned}$$

$$S = \int d^4x \left[-\frac{1}{2} \partial_\mu X \partial^\mu X - \tilde{V}(X) \right]$$

(The th. in terms of X is a perfectly good th.)

X has $(\text{mass})^2 = 2\mu^2 > 0$ \Rightarrow a well defined perturbation theory
 regular field (not a tachyonic field)

X^3 & X^4 are interactions

However this does not have a manifest symmetry of the correlation functions

The $\phi \rightarrow -\phi$ symmetry is spontaneously broken.

(Even though the original action had this sym., the physics that we get out of it won't have this sym. — our correl. fun. of X are the physical quantities & they don't have this sym.)

e.g. if ϕ is
3 or 4 w.r.t.
of X won't
see this sym.

Reason for symmetry breaking:-

① Path integral formalism:-

For a conventional theory of +ve $(\text{mass})^2$ we use a boundary

condition $\phi \rightarrow 0$ as $x \rightarrow \infty$

(otherwise we get various kinds of boundary terms which are hard to deal with — well defined by b.c. so make the path int. For an infinite vol. sys., what b.c. we have put on the fields don't affect the physics)

In a spontaneously broken symmetric theory, the natural b.c. is

$$X \rightarrow 0 \text{ as } x \rightarrow \infty$$

$$\phi \rightarrow \sqrt{u^2/2} \text{ as } x \rightarrow \infty \quad \left| \begin{array}{l} \phi \rightarrow -\sqrt{u^2/2} \\ \text{is also possible} \end{array} \right.$$

↑ symmetry
(relates these 2 b.c.)

[Physically both b.c. are equivalent]

- In path int., the b.c. should also be inv.
- But there is a one-to-one map betw. the corrly fns. for the 2 b.c.
- But the corrly fns. by themselves don't exhibit this sym. as ~~the~~ b.c. are diff. & path int. will give diff. results for diff. b.c.]

Canonical formalism

$$\langle \phi(x_1) \dots \phi(x_n) \rangle$$

$$= \langle \omega | T(\hat{\phi}(x_1) \dots \hat{\phi}(x_n)) | \omega \rangle$$

(the symmetry) $\phi(x) \rightarrow -\phi(x)$ is implemented by a unitary operator \hat{U} :-

$$\hat{U} \hat{\phi}(x) \hat{U}^{-1} = -\hat{\phi}(x)$$

$$\langle \omega | \hat{U} (T(\hat{\phi}(x_1) \dots \hat{\phi}(x_n))) \hat{U}^{-1} | \omega \rangle$$

$$= \langle \omega | T \left(\underbrace{\hat{U} \hat{\phi}(x_1) \hat{U}^{-1}}_{-\hat{\phi}(x_1)} \underbrace{\hat{U} \hat{\phi}(x_2) \hat{U}^{-1}}_{-\hat{\phi}(x_2)} \dots \right) | \omega \rangle$$

\hat{U} is a discrete sym. - for a continuous sym. you would have continuous parametr

$$\Rightarrow \langle \alpha | \hat{U}^{\dagger} (\hat{\phi}(x_1) \dots \hat{\phi}(x_n)) \hat{U} | \alpha \rangle$$

$$= (-1)^n \langle \alpha | T(\hat{\phi}(x_1) \dots \hat{\phi}(x_n)) | \alpha \rangle$$

Now, $\hat{U}^{\dagger} \hat{H} \hat{U}^{-1} = \hat{H}$ \rightarrow since \hat{U} is a symmetry.

$$\therefore \hat{U}^{\dagger} \hat{H} \hat{U}^{-1} | \alpha \rangle = \hat{H} | \alpha \rangle = E_{\alpha} | \alpha \rangle$$

↑
vacuum
energy

$$\Rightarrow \hat{H} \hat{U}^{-1} | \alpha \rangle = E_{\alpha} \hat{U}^{-1} | \alpha \rangle$$

In other words, $\hat{U}^{-1} | \alpha \rangle$ also has energy E_{α} (same as that of $| \alpha \rangle$)

Now there two possibilities :-

$$\textcircled{1} \quad \hat{U}^{-1} | \alpha \rangle = | \alpha \rangle \Rightarrow \langle \alpha | \hat{U} = \langle \alpha |$$

(In this case we are done)

$$(\text{then}) \Rightarrow \langle \alpha | T(\hat{\phi}(x_1) \dots \hat{\phi}(x_n)) | \alpha \rangle$$

$$= (-1)^n \langle \alpha | T(\hat{\phi}(x_1) \dots \hat{\phi}(x_n)) | \alpha \rangle$$

(This is the case when sym. is unbroken - the action of sym. frs. leaves the vacuum unchanged & we get the same sym. in the corrly frs. bcs of this)

$$\textcircled{2} \quad \hat{U}^{-1} | \alpha \rangle = | \alpha' \rangle \rightarrow \text{different state}$$

$| \alpha' \rangle$ can also be called a vacuum bcs it has the same energy $E_{\alpha'}$ as $| \alpha \rangle$ — degeneracy in vacuum state - typically

Sym. trs. will take you from one to the other
 — mix ... — doesn't leave the vacuum invariant)

Why is it not the gr. state is a superposition
 of a wave func centred around the 2 minima as
 in H.O.?

→ On QFT, we have an α no. H.O.'s

$$\left\{ \begin{array}{l} \langle \psi_1 | \hat{\pi}_1 | \psi_2 \rangle = \text{a small no.} \\ \text{perturbative ground states} \\ \hat{\pi}_1 = \begin{pmatrix} \langle \psi_1 | \hat{\pi}_1 | \psi_1 \rangle & \langle \psi_1 | \hat{\pi}_1 | \psi_2 \rangle \\ \langle \psi_2 | \hat{\pi}_1 | \psi_1 \rangle & \langle \psi_2 | \hat{\pi}_1 | \psi_2 \rangle \end{pmatrix} \end{array} \right.$$

1st order pert.

On QFT, these are raised to infinite power
 → overlap is zero → true ground state is
 not a superposition of the two — either $|\psi_1\rangle$
 or $|\psi_2\rangle$

$$\hat{\pi}_1 = \hat{\pi}_0 + \hat{\pi}_1$$

$$\langle -\psi_0 | \hat{\pi}_1 | -\psi_0' \rangle = 0$$

↑ pert. gr. states

becos it will involve prod. of α
 no. of terms each
 of which is less
 than 1

In path int., b.c. fixing means as if freezing

~~the has occurred~~ occurred at the boundary
analogy → stat. mech.
 $m > 0$ → like in paramagnetic state & system
 doesn't know about boundary

$\mu < 0$ → .. " ferrromag. phase & b.c.
 makes a diff. — you have frozen
 spins at the boundary & this will
 align the rest of the spins

Sym. is broken by dynamics though the Lag. has the sym. — spontaneous sym. breaking
 ∵ we have moved to a new vacuum — imposed certain b.c. — now if we construct corrly. of ϕ after calculating these corrly. of ϕ also don't obey the sym. — bcs the b.c. are diff.
 Exp. we see only X -particles

a, at here meaning only if you can apply pert. th.
 — otherwise meaningless

$$\text{Note} \rightarrow \langle r_2 | \phi(r) | r_2 \rangle \neq 0$$

$$\text{while } \langle r_1 | \phi(r) | r_2 \rangle = 0$$

Spontaneous breaking of continuous symmetries

$$S = \int d^4x \left[-\eta^{\mu\nu} \partial_\mu \phi^* \partial_\nu \phi - V(\phi) \right]$$

$$\text{where } V(\phi) = -\frac{1}{2} \mu^2 \phi^* \phi + \frac{\lambda}{4} (\phi^* \phi)^2$$

$$\phi = \frac{1}{\sqrt{2}} (\phi_1 + i\phi_2) \text{ is complex}$$

$$S = \int d^4x \left[-\frac{1}{2} \eta^{\mu\nu} (\partial_\mu \phi_1 \partial_\nu \phi_1 + \partial_\mu \phi_2 \partial_\nu \phi_2) - V \right]$$

$$\text{where } V = -\frac{1}{2} \mu^2 (\phi_1^2 + \phi_2^2) + \frac{\lambda}{4} (\phi_1^2 + \phi_2^2)$$

Symmetry of S

$$\phi \rightarrow e^{i\alpha} \phi, \quad \phi^* \rightarrow e^{-i\alpha} \phi^*$$

or in terms of real fields

$$\begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix} \rightarrow \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix}$$

(i) Φ potential has an axial sym. \rightarrow corr. to rotation in xy -plane)

$$V(\phi) = -\mu^2 |\phi|^2 + \lambda |\phi|^4$$

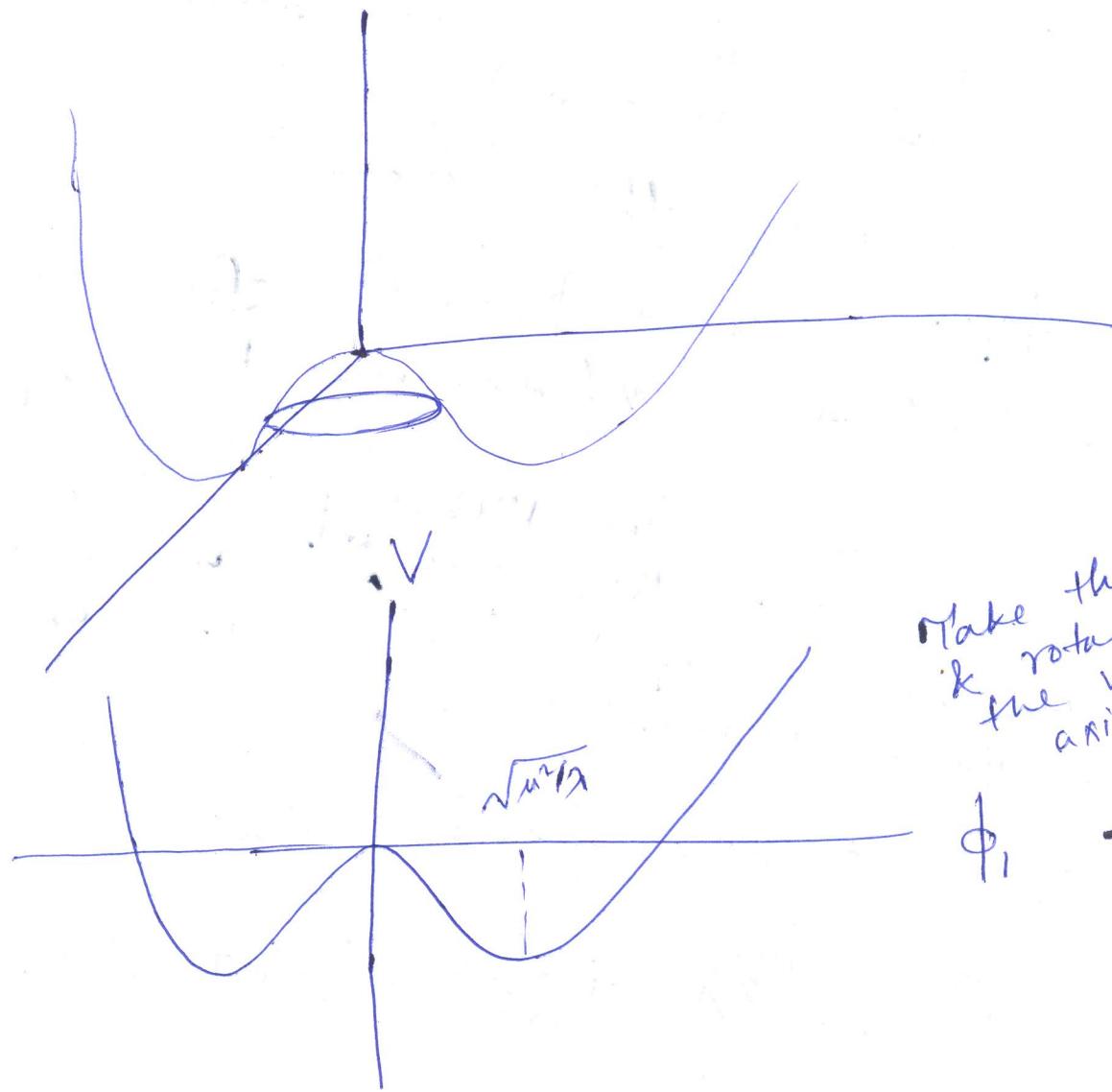
$$\frac{\partial V}{\partial |\phi|} = -2\mu^2 |\phi| + 4\lambda |\phi|^3 \Rightarrow |\phi| = \sqrt{\frac{\mu^2}{2\lambda}}$$

$$\therefore \phi = \sqrt{\frac{\mu^2}{2\lambda}} e^{i\theta}, \text{ where } \theta \text{ arbitrary}$$

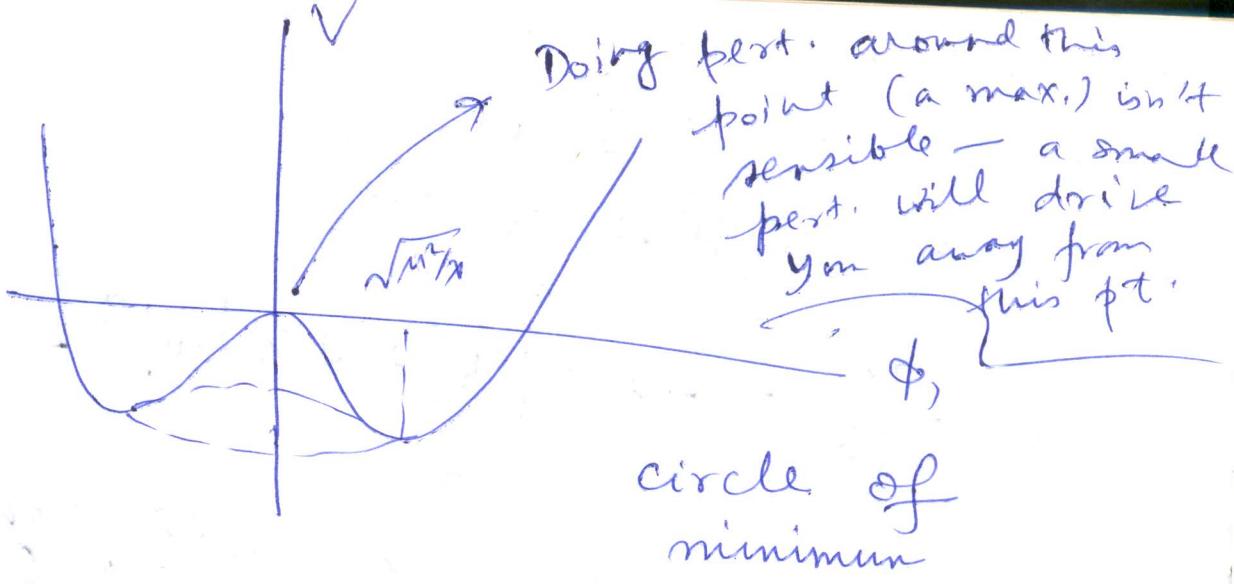
(Minimize w.r.t. the real field or the complex field
 \rightarrow will get the same ans.)

$$\phi_1 = \sqrt{\frac{\mu^2}{2\lambda}} \cos \theta$$

$$\phi_2 = \sqrt{\frac{\mu^2}{2\lambda}} \sin \theta$$



Make this figure & rotate about the vertical axis by 360°
 \rightarrow generate in the full $\phi_1 + \phi_2$ phase



$$\phi_1 = \sqrt{m\gamma_2} \cos \theta \quad \text{and} \quad \phi_2 = \sqrt{m\gamma_2} \sin \theta \quad \begin{matrix} \nearrow \\ \text{labels a circle of} \\ \text{minimum} \end{matrix}$$

(Physically it doesn't matter which pt. on the circle we choose for pert. — we can map this corresp. to others about diff. pts. of minima)

Let us choose the point

$$\phi_1 = \sqrt{m\gamma_2}, \phi_2 = 0 \quad \text{for the perturbation expansion } \phi = \sqrt{\frac{m}{2}} \chi$$

(There is no overlap betw. the Hilbert spaces bcos of infinite vol. — otherwise the ^{correct} vacuum would be a superposition of the diff. vacuum states — Hilbert spaces are distinct)

Define a new field χ through

$$\phi = \sqrt{\frac{m}{2}} \chi + \frac{1}{\sqrt{2}} (\chi_1 + i\chi_2)$$

$$S = \int d^4x e \left[-\frac{1}{2} \eta^{\mu\nu} \partial_\mu X_1 \partial_\nu X_1 + \frac{1}{2} \eta^{\mu\nu} \partial_\mu X_2 \partial_\nu X_2 - \tilde{V}(X_1, X_2) \right]$$

where $\tilde{V}[X_1, X_2] = V[\phi] = -\lambda/4 \left(\frac{\mu^2}{2}\right)^2 + \mu^2 X_1^2 + \text{cubic. \& higher order terms}$

Ex. Prove this

X_1 has $(\text{mass})^2 = 2\mu^2$

X_2 has $(\text{mass})^2 = 0$

GOLDSTONE BOSONS

this feature is generic in a th. of spontaneously broken sym.

Pot. is flat along the angular dirn. — 2nd deriv. of pot. along this dirn. is zero — In orthogonal dirn., pot. changes — Shifting X_1 means moving in orthogonal dirn. & we get a $(\text{mass})^2$ — X_2 corr. to a massless particle & this in turn corr. to moving in the angular dirn.

sym. trs. parameter was a bosonic coordinate

parameter → these fluctuations are bosonic & the corr. massless particles are bosons.

Spontaneously broken fermionic sym. will give fermions.

$$\phi = (\sqrt{m_{22}} + \delta) e^{i\theta}$$

\rightarrow It becomes clear that $V(\phi)$ is indep. of θ

\rightarrow take ϕ & θ as our fields.

To the leading order, X_2 is the same as θ up to a normalisation (bcos X_2 is the imag. part of ϕ)

Temperature can restore spontaneously broken sym.
 \rightarrow e.g. in ferromagnet.