

Statistical Mechanics Exam: Part I

1. Consider a cylinder of radius a , kept vertically, containing N molecules of a monatomic gas. m is the mass of each atom, g is the acceleration due to gravity, and T is the temperature of the gas. The upper end of the cylinder is closed with a piston of weight w . In the equilibrium situation the piston is at a height h above the lower end of the cylinder. The height h is large so that the effect of the gravitational pull on the atoms cannot be ignored. The temperature is sufficiently large so that we can use classical statistical mechanics.
 - a) Find an expression for the free energy F of the system as a function of a, N, m, g, h .
 - b) Using the result of part a), and the equation $dF = -dW - SdT$ where dW is the work done by the system, and S is the entropy of the system, calculate the net force exerted by the gas on the piston, i.e. calculate the weight w in terms of a, N, m, g, h .
 - c) Calculate the equilibrium density of atoms at a height x above the bottom of the cylinder.
 - d) Using the result of part c), calculate the pressure at the top of the cylinder and the force exerted by the gas on the piston. Compare with the result of part a).
2. Consider a two dimensional ideal monatomic Bose gas containing N atoms each of mass m , confined in an area A with the geometry of a square, and periodic boundary condition in both directions. Find the temperature T at which the fraction f (f is small compared to 1 but fN is large compared to 1) of the total number of particles will be in the ground state. T should be determined in terms of N, A and m . (Note that you need to do this calculation for large but finite A .)
3. Consider a lattice containing two rows of sites as follows

x	x	x	x	x	x	x	x	...	x	x	x	x	x	x	x	x
x	x	x	x	x	x	x	x	...	x	x	x	x	x	x	x	x

Each row contains $2N$ sites, with the $(2N + 1)$ th site identified with the first site. To the i th site in the upper row we associate a variable

s_i , and to the i th site in the lower row we associate a variable σ_i . s_i and σ_i take values ± 1 . The total energy of the system for a given configuration of spins is given by:

$$E(\{s_i\}, \{\sigma_i\}) = -J \sum_{i=1}^{2N} (\sigma_i s_{i+1} + s_i \sigma_{i+1})$$

Find an expression for the free energy and the specific heat of this system.

~~1/2 hole~~

Fermions with spins

→ $(2s+1)$ degenerate states for every \vec{p} .
[in absence of mag. field]

$$\ln \Omega = (2s+1) \frac{V}{h^3} \int d^3p \ln (1 + 2e^{-\beta e(\vec{p})})$$

$$N = (2s+1) \frac{V}{h^3} \int d^3p \frac{2e^{-\beta e(\vec{p})}}{1 + 2e^{-\beta e(\vec{p})}}$$

$$\frac{PV}{NkT} = V/N \frac{\partial}{\partial N} (\ln \Omega) = \frac{F_1(z)}{F_2(z)} > \text{same functions (as before)}$$

$$\rightarrow \frac{N}{2s+1} = \frac{V}{h^3} \int d^3p \frac{2e^{-\beta e(\vec{p})}}{1 + 2e^{-\beta e(\vec{p})}}$$

→ same eqns as in the earlier case with N replaced by $\frac{N}{2s+1}$.

Ideal Bose gas

$$\ln \Omega = - \sum_{\vec{p}} \ln (1 - 2e^{-\beta e(\vec{p})})$$

$$\bar{N} = \sum_{\vec{p}} \frac{2e^{-\beta e(\vec{p})}}{1 - 2e^{-\beta e(\vec{p})}}, \quad \bar{n}(\vec{p}) = \frac{2e^{-\beta e(\vec{p})}}{1 - 2e^{-\beta e(\vec{p})}}$$

$$\bar{n}(\vec{p}) \geq 0 \text{ for every } \vec{p}$$

free up to here

$$z \leq 1$$

[why is this happening? →

$$B \cos \left(\sum_k \frac{2e^{-\beta e(\vec{p})}}{1 - 2e^{-\beta e(\vec{p})}} \right)^k = \frac{1}{1 - 2e^{-\beta e(\vec{p})}}$$

only when

$$z \leq 1$$

$$N = \sum_{\vec{p}} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}}$$

Used to solve for z in terms of N

(Prop. of the R.H.S. !)

$$\sum_{\vec{p}} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}} = \sum_{\vec{p}} \frac{1}{z^{-1} e^{\beta e(\vec{p})} - 1}$$

\downarrow
increases as
 z increases

As $z \rightarrow 0$

$$\sum_{\vec{p}} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}} \xrightarrow{z \rightarrow 0} 0$$

for small z , we make the approximation
 $z \sum_{\vec{p}} e^{-\beta e(\vec{p})}$
this thing is finite
so neglecting (cos it is sufficiently convergent for large \vec{p})
the sum is justified (we aren't making any div. by \vec{p})
we aren't making any div. by \vec{p} this approximation

As $z \rightarrow 1$

$$\sum_{\vec{p}} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}} = \frac{z}{1-z} + \sum_{\vec{p} \neq 0} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}}$$

free bosons, & so
 $e(\vec{p}) = \vec{p}^2 / 2m$

\rightarrow as $z \rightarrow 1$

$$\Rightarrow \bar{N} = \sum_{\vec{p}} \frac{z e^{-\beta e(\vec{p})}}{1 - z e^{-\beta e(\vec{p})}}$$

in the range $0 < z < 1$
has a unique soln. for
any $0 < \bar{N} < \infty$

An overall shift in energy can be absorbed in the redefn. of z

bcz \bar{N} is monotonically ↑ fn. of z .

Take the $V \rightarrow \infty$ limit & replace the sum by integral.

$$\ln Q = -\frac{V}{k^3} \int d^3 p \ln \left(1 - z e^{-\beta \vec{p}^2/2m} \right)$$

$$= -\frac{V}{k^3} 4\pi \int p^2 dp \ln \left(1 - z e^{-\beta p^2/2m} \right)$$

$$p = |\vec{p}|$$

$$\ln Q = -\frac{4\pi V}{k^3} (2m k T)^{3/2} \int_0^\infty u^2 du \ln \left(1 - z e^{-u^2} \right)$$

$$p = \sqrt{\frac{2m}{\beta}} \cdot u$$

$$= \frac{4\pi V}{k^3} (2m k T)^{3/2} F_1(z)$$

$$\hookrightarrow - \int_0^\infty u^2 du \ln \left(1 - z e^{-u^2} \right)$$

$$\bar{N} = \frac{V}{k^3} \int d^3 p \frac{z e^{-\beta \vec{p}^2/2m}}{1 - z e^{-\beta \vec{p}^2/2m}}$$

$$= \frac{4\pi V}{k^3} (2m k T)^{3/2} F_2(z)$$

$$\text{where } F_2(z) = \int_0^\infty u^2 du \frac{z e^{-\beta u^2}}{1 - z e^{-\beta u^2}}$$

Need to solve

$$\boxed{\bar{N} = \frac{4\pi V}{k^3} (2m k T)^{3/2} F_2(z)}$$

$$F_2(z) = \int_0^\infty u^2 du \frac{1}{z^{-1} e^{u^2} - 1}$$

increases
as
 z increases

(It was seen in the discrete version - so no surprise that we see the same behavior in the continuum version.)

In order to get a good stat. ensemble for bosons, we have to restrict z to $0 < z < 1$ & for fermions $0 < z < \infty$

$e^{\beta u N} \rightarrow$ weight factor
Allowed wt. factor as long as the sum is convergent
- not physically --- is non-conv.

μ is an auxiliary variable - μ is a trick we have introduced

z is an intermediate contraction to solve N

~~For bosons~~
Grand canonical ensemble won't work when you find the sum is non-conv.
→ maybe use some other ensemble

But here we see we can use grand can. for N
 $\infty < N < \infty$

$$\text{As } z \rightarrow 0, F_2(z) \simeq \int_0^\infty u^2 du z e^{-u^2}$$

$$\simeq z \times \text{finite integral}$$

$$\rightarrow 0 \text{ as } z \rightarrow 0$$

As $z \rightarrow 1$

$$F_2(z) \rightarrow \int_0^\infty \frac{u^2 du}{e^{u^2} - 1}$$

2 things to check:
whether it is div. for
large u region &
small u region

for small u ,

$$\int du u^2 \frac{1}{1+u^2-1} \sim \int du \rightarrow \text{finite}$$

As $u \rightarrow 0$,
there is a
pole of the
integrand

$\frac{1}{e^{u^2}-1} \sim \frac{1}{u^2}$ as $u \rightarrow 0$
even if the integrand may
div., the integral may
not diverge

$$\text{e.g. } \int_0^1 \frac{dx}{x^2 + x^4} \sim \int_0^1 \frac{dx}{x^2} = \frac{1}{2} x^{-1} \Big|_0^1 = \frac{1}{2}$$

Without taking a
radial cond. (i.e., taking
 $dz = dx + iy$) we could have
used $\int_{\gamma(t^2+y^2+z^2)=1} dz$

to give the
same result

$$\int_0^1 + \int_1^\infty$$

$$\int_0^1 \frac{u^2 du}{e^{u^2}-1} = \int_0^1 \frac{u^2 du}{u^2 + u^4 + \dots}$$

$$= \int_0^1 u^2 du \frac{1}{u^2} \{ 1 + u^2 + \dots \}^{-1}$$

$$\int_0^\infty = \int_0^1 + \int_1^\infty$$

↓ ↓
finite finite

Suppose $F_2(z) \rightarrow K$ as $z \rightarrow 1$

↓
calculate
constant

$$K = \int u^2 du \frac{e^{-u^2}}{1-e^{-u^2}}$$

$0 < F_2(z) < K$ if $0 < z < 1$

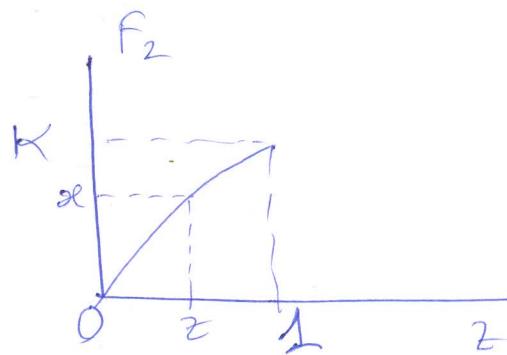
$$N = \frac{4\pi V}{h^3} (2m\kappa T)^{3/2} F_2(z)$$

$$\text{if } \bar{N} < \frac{4\pi V}{h^3} (2m\kappa T)^{3/2} K$$

then this eqn has a unique solution
in the range $0 < z < 1$

$$z = \frac{h^3}{\bar{N}} \frac{1}{4\pi V} (2m\kappa T)^{3/2}$$

$$z = F_2(z)$$



$$\text{if } \begin{cases} z > K \\ \bar{N} > \end{cases} \Rightarrow \frac{4\pi V}{h^3} (2m\kappa T)^{3/2} K$$

\Leftrightarrow same as
then there is no soln to

$$\bar{N} = \frac{4\pi V}{h^3} (2m\kappa T)^{3/2} F_2(z)$$

in the range $0 < z \leq 1$.

Go back to the discrete version

$$\bar{N} = \sum_{\vec{P}} \frac{2e^{-\beta e(\vec{P})}}{1 - 2e^{-\beta e(\vec{P})}}$$

$$\sqrt{\frac{1}{h^3}} \int d^3p \quad \text{requires that} \quad \frac{2e^{-\beta e(\vec{P})}}{1 - 2e^{-\beta e(\vec{P})}}$$

is smooth

We can go from one pt. in the lattice to another, if it doesn't jump by a large amt.

$$\text{if } z < 1 \text{ then } \frac{2e^{-\beta e(\vec{P})}}{1 - 2e^{-\beta e(\vec{P})}} \text{ is finite}$$

sth has gone wrong in going from discrete to continuum version
∴ in discrete version we had a soln for every value of \bar{N} , but in continuum version we find this is not true

If $z \approx 1$ then the summand is not smooth at $e(\tilde{f}) \approx 0$.

Take $\frac{1}{g}$

Suppose you are starting from ~~far~~^{far} case even if $\epsilon = \epsilon_0$, $f = 2\epsilon$

$\frac{1}{\epsilon}, \frac{1}{2\epsilon}$ differ by large amt.

↓ can't write this as an integral

$\frac{1}{g_{\infty}}$ & $\frac{1}{\epsilon}$ is enormous

$$e(\tilde{f}^{(g_0)}, 1) = \frac{1}{2m} \left(\frac{2\pi N t}{\epsilon} \right)^2$$

↓ also $\rightarrow 0$ in the thermodynamic limit — so you can't say that $e(\tilde{f}^{(g_0)})$ is separate & treat the rest as integral) problem is not only at $\epsilon \approx 0$, the problem is at small ϵ , in principle

Lesson:— As $z \rightarrow 1$ we cannot use the naive thermodynamic limit of replacing ~~the sum~~ by $\frac{V}{h^3} \int d^3 p$

$$g_f \bar{N} \leq \frac{4\pi V}{h^3} (2mkt)^{3/2} K$$

$$\text{then } \bar{N} = \frac{4\pi V}{h^3} (2mkt)^{3/2} F_2(z)$$

has a unique solution for $0 < z < 1$.

↓ thermodynamic limit has no subtlety

$z=1$ has a problem, so we don't take $\bar{N} \leq \dots$

self-consistent way of det. z

↓ assume the thermodyn. limit, find $z = \dots$ & then check see that the thermodyn. limit is valid

$$\Rightarrow T \geq \left(\frac{\bar{N}}{4\pi V K} \right)^{1/3} \frac{h^2}{2mk} \equiv T_c$$

↓ critical temperature

We can use this to study high temperature limit.

$\beta \rightarrow 0$ ($T \rightarrow \infty$) at fixed N/k .

$$\frac{N}{T} = \frac{4\pi}{h} (2m k T)^{3/2} F_2(z)$$

If $T \rightarrow \infty$ at fixed N/k , then

$$F_2(z) \rightarrow 0$$

$$\downarrow z \rightarrow 0$$

$$F_1(z) = - \int_0^\infty u^2 du \ln(1 - ze^{-u^2})$$

$$\approx - \int_0^\infty u^2 du \left\{ 1 - ze^{-u^2} - \frac{1}{2} z^2 e^{-2u^2} - \dots \right\}$$

$$\approx \sqrt{\frac{\pi}{2}} \left(2 + \frac{1}{4\sqrt{2}} z^2 + \dots \right)$$

$$F_2(z) = \int_0^\infty u^2 du \frac{ze^{-u^2}}{1 - ze^{-u^2}}$$

$$= \int_0^\infty u^2 du \left(2e^{-u^2} + z^2 e^{-2u^2} + \dots \right)$$

$$\approx \frac{\sqrt{\frac{\pi}{2}}}{4} \left(2 + \frac{1}{2\sqrt{2}} z^2 + \dots \right)$$

same integral as
for fermions
except for
the '-' sign
before ze^{-u^2}

$$\frac{PV}{NkT} = \frac{F_1(z)}{F_2(z)} = \frac{\left(1 + \frac{1}{4\sqrt{2}} z^2 + \dots \right)}{\left(1 + \frac{1}{2\sqrt{2}} z^2 + \dots \right)}$$

$$\approx \left(1 - \frac{1}{4\sqrt{2}} z^2 + \dots \right)$$

$$\frac{N}{V} = \frac{4\pi}{h^3} (2m\kappa T)^{3/2} \sqrt{\pi}/g_2 \left(1 + \frac{2}{252} + \dots \right)$$

Leading soln :-

$$Z = \frac{N}{V} \frac{h^3}{\pi \sqrt{\pi}} (2m\kappa T)^{-3/2} + \dots$$

\therefore we get

$$\frac{PV}{NkT} = \left\{ 1 - \frac{1}{4\sqrt{2}} \frac{\sqrt{\pi}}{V} \frac{h^3}{(\pi)^{3/2}} \frac{1}{(2m\kappa T)^{3/2}} + \dots \right\}$$

classical ideal gas answer

again illustrates the fact that classical stat. is somewhere in between Bose & Fermi.
Bose \rightarrow more states available than for classical

Case for $T < T_c$:-

$$N = \sum_{E(P)} \frac{2e^{-\beta e(P)}}{1 - 2e^{-\beta e(P)}}$$

We need to take $2 \rightarrow 1$
as $N \rightarrow \infty$.

When the thermodyn. limit breaks down

(ϵ close to 1 as you are taking the thermodyn. limit)

$$\bar{N} = \sum_{\substack{P \\ \epsilon(P) < \eta}} \frac{2e^{-\beta e(P)}}{1 - 2e^{-\beta e(P)}} + \sum_{\substack{P \\ \epsilon(P) > \eta}} \frac{2e^{-\beta e(P)}}{1 - 2e^{-\beta e(P)}}$$

\downarrow

I_1 I_2

Now η ! A fixed number
take $V \rightarrow \infty$ limit.
(In this case we know that somehow $Z \rightarrow 1$)

(in this limit) 2nd term

$$I_2 = \frac{N}{h^3} \int d^3 p \frac{2e^{-\beta E/2m}}{1 - 2e^{-\beta E/2m}}$$

$$= G(\eta, \beta, \varepsilon) \rightarrow \begin{array}{l} \# \text{ of particles with} \\ \text{energy } \geq \eta \end{array}$$

some function

$$I_2 \xrightarrow{\eta \rightarrow 0} \frac{4\pi V}{h^3} (2mkT)^{3/2} K$$

η -dep. of G
is the imp. part

of particles
with energy $\geq \eta$
as $\eta \rightarrow 0$

$$I_1 = \sum_{E(\vec{p}) < \eta} \frac{2e^{-\beta E(\vec{p})}}{1 - 2e^{-\beta E(\vec{p})}}$$

$E(\vec{p}) < \eta$ = # of particles
with energy $< \eta$

$$= N - I_2$$

$$\eta \rightarrow 0 \quad \left\{ 1 - \frac{4\pi V}{N h^3} (2mkT)^{3/2} K \right\}$$

some fraction > 0

$$N \geq \frac{4\pi V}{h^3} (2mkT)^{3/2} K$$

of particles with energy η in the $\eta \rightarrow 0$ limit
is a finite fraction of the total no.
of particles.

the fact that we cannot account for the
particles with energy close to zero shows that
thermodyn. breaks down for them

5/11/06

Ideal Bose gas

$$N = \sum_{\vec{p}} \frac{2e^{-\beta e(\vec{p})}}{1 - 2e^{-\beta e(\vec{p})}} = \sum_{\vec{p}} \frac{1}{z^{-1}e^{\beta e(\vec{p})} - 1} = h(z)$$

$\underbrace{z^{-1}e^{\beta e(\vec{p})} - 1}_{n(\vec{p})}$

- 1) $n(\vec{p})$ for each \vec{p} is a monotone increasing function of z for $0 < z < 1$
- 2) $\Rightarrow h(z)$ is a monotone increasing function of z .
- 3) $h(z) \rightarrow 0$ as $z \rightarrow 0$
- 4) $h(z) \rightarrow \infty$ as $\cancel{z \rightarrow 1}$.
- 5) Replacing (the summation by an integration for large V)

$\sum_{\vec{p}}$ by $\frac{V}{h^3} \int d^3p$ for large V breaks down when $z \approx 1$, $e(\vec{p}) \approx 0$.

$$\bar{N} = \sum_{\substack{\vec{p} \\ e(\vec{p}) < \epsilon}} \frac{1}{z^{-1}e^{\beta e(\vec{p})} - 1} + \sum_{\substack{\vec{p} \\ e(\vec{p}) \geq \epsilon}} \frac{1}{z^{-1}e^{\beta e(\vec{p})} - 1}$$

\downarrow
 $u \geq \beta \epsilon$
 $V \rightarrow \infty$

ϵ : fixed energy.

$$\lim_{\epsilon \rightarrow 0} H_2(z, \epsilon) = F_2(z)$$

$$H_2(\infty, \epsilon) \leq F_2(z)$$

$$\int_0^\infty u^2 du \frac{1}{z^{-1}e^{u^2} - 1}$$

$$\frac{4\pi V}{h^3} (2m\kappa T)^{3h} \quad \cancel{H_2(\beta)}$$

$$H_2(z, \epsilon) = \int_{\sqrt{\beta \epsilon}}^\infty u^2 du \frac{1}{z^{-1}e^{u^2} - 1}$$

$$F_2(z) \rightarrow 0 \text{ as } z \rightarrow 0$$

$\rightarrow K_0$ as $z \rightarrow 1$

$$\int_0^\infty u^2 du \frac{1}{e^{u^2} - 1} = \text{finite constant}$$

$$\frac{\bar{N}}{V} = \frac{1}{V} \sum_{\substack{p \\ e(p) < \epsilon}} \frac{1}{z^{-1} e^{\beta e(p)} - 1}$$

$$+ \frac{4\pi}{h^3} (2m\kappa T)^{3/2} H_2(z, \epsilon)$$

$$\lim_{\epsilon \rightarrow 0} \quad \lim_{V \rightarrow \infty}$$

$$\frac{\bar{N}}{V} = \lim_{\epsilon \rightarrow 0} \lim_{V \rightarrow \infty} \frac{1}{V} \sum_{\substack{p \\ e(p) < \epsilon}} \frac{1}{z^{-1} e^{\beta e(p)} - 1}$$

$$+ \frac{4\pi}{h^3} (2m\kappa T)^{3/2} F_2(z)$$

$\frac{\bar{N}}{V}$ is a quantity that we're given

$$\text{Define: } g = \frac{\bar{N}}{V}$$

Define T_c through

$$g = \frac{4\pi}{h^3} (2m\kappa T_c)^{3/2} K_0 \rightarrow \text{defn. of } T_c$$

for a given g , we define T_c by this eqn. given
 T_c isn't an intrinsic prop. of the material it depends upon g

$$\frac{4\pi}{h^3} (2m\kappa T_c)^{3/2} K_0 - \frac{4\pi}{h^3} (2m\kappa T)^{3/2} F_2(z) \\ = \lim_{\epsilon \rightarrow 0} \lim_{V \rightarrow \infty} \frac{1}{V} \sum_{\substack{p \\ e(p) < \epsilon}} \frac{1}{z^{-1} e^{\beta e(p)} - 1}$$

If $T < T_c$, then L.H.S. is a positive number since $F_2(z) \leq K_0$.

Hence, right hand side $\neq 0$ \rightarrow this can happen only if $z \rightarrow 1$ as $V \rightarrow \infty$

$$\Rightarrow F_2(z) \rightarrow K_0$$

But if $z < 1$, we can replace \sum by \int & the upper limit being ϵ , it goes to zero as $\epsilon \rightarrow 0$

$$\frac{4\pi}{\hbar^3} k_0 (2m\hbar)^{\frac{3k}{2}} \left\{ T_c^{\frac{3k}{2}} - T^{\frac{3k}{2}} \right\}$$

$$= \lim_{E \rightarrow 0} \lim_{V \rightarrow \infty} \frac{1}{V} \sum_{\substack{\vec{p} \\ e(\vec{p}) < E}} \frac{1}{z^{-1} e^{\beta e(\vec{p})} - 1}$$

R.H.S. = $\frac{1}{V} \times$ Total no. of particles with energy $< E$

for $T(T_c)$, the total number of particles with energy $< E$

$$= V \times \text{some fixed no.}$$

$$= \frac{N}{V} \times \text{some fixed no.}$$

$$\boxed{P = N/V \text{ fixed}}$$

Fraction of particles with energy $< E$
 $=$ a fixed number even as $E \rightarrow 0$

This is not what we expect, we expect that if energy E goes to zero as $E \rightarrow 0$

Q) How are these particles distributed among the energy levels $< E$?

Consider a box of lengths

$$(L_1, L_2, L_3)$$

$$\rho = \frac{4\pi^2 \hbar^2}{2m} \left(\frac{n_1^2}{L_1^2} + \frac{n_2^2}{L_2^2} + \frac{n_3^2}{L_3^2} \right)$$

$$= \frac{2\pi^2 \hbar^2}{m L^2} (n_1^2 + n_2^2 + n_3^2)$$

$$L_1 = L_2 = L_3 = L$$

for simplicity

$$\frac{1}{V} \sum_{\substack{\vec{p} \\ e(\vec{p}) < E}} \frac{1}{z^{-1} e^{\beta e(\vec{p})} - 1}$$

$$\begin{aligned} \text{R.H.S.} &= \frac{1}{V} \sum_{\substack{\vec{p} \\ e(\vec{p}) < E}} \frac{1}{z^{-1} e^{\beta e(\vec{p})} - 1} + \sum_{\substack{\vec{p} \neq 0 \\ e(\vec{p}) < E}} \frac{1}{z^{-1} e^{\beta e(\vec{p})} - 1} \\ &= \frac{1}{V} \frac{1}{z^{-1} - 1} + I_2 \end{aligned}$$

We can set $z=1$ in the LHS, not in the RHS bcos $z \rightarrow 1$ only as $V \rightarrow \infty$. In the LHS the V -dependence has dropped out & so we could put $z=1$ there.

We can set $z=1$ only in those terms where $V \rightarrow \infty$ limit is smooth.

First $V \rightarrow \infty$ has been taken & then $E \rightarrow 0$; so E is much greater than energy spacings; we have already made the spacings approach zero.

We will show

$$\lim_{\epsilon \rightarrow 0} \lim_{V \rightarrow \infty} I_2 = 0$$

\Rightarrow All the contribution must come from I_1

\Rightarrow A finite fraction of the particles will be in $\vec{p} = 0$ state.

Each term in I_2 is positive.

$$I_2 \leq \frac{1}{V} \sum_{\vec{p} \neq 0} \frac{1}{e^{\beta e(\vec{p})} - 1}$$

$$e(\vec{p}) \leq \epsilon$$

(\because it was a monotone \uparrow fn. of ϵ & if we put $\epsilon=1$, that is maximum)

$$e^{\beta \epsilon} > 1 + \beta \epsilon$$

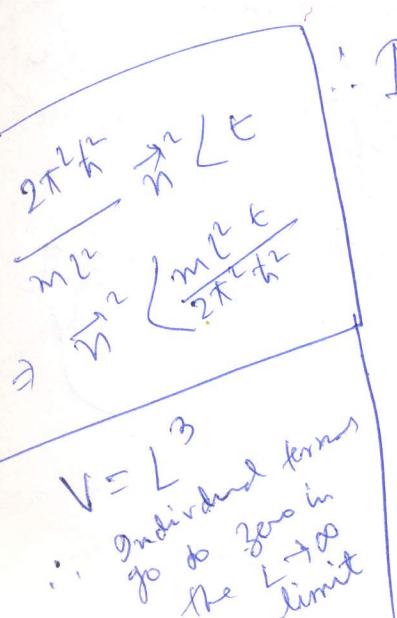
$$\therefore I_2 \leq \frac{1}{V} \sum_{\vec{p} \neq 0} \frac{1}{1 + \beta e(\vec{p}) - 1}$$

$$= \frac{1}{PV} \sum_{\vec{n} \neq 0} \frac{1}{\frac{4\pi^2 h^2}{2mL^2} (n_1^2 + n_2^2 + n_3^2)}$$

$$\epsilon \leq \epsilon$$

$$\therefore I_2 \leq \frac{1}{PV} \frac{mL^2}{2\pi^2 h^2} \sum_{\vec{n} \neq 0} \frac{1}{n_1^2 + n_2^2 + n_3^2}$$

$$\vec{n}^2 < \frac{mL^2}{2\pi^2 h^2}$$



When $\epsilon \sim \frac{1}{L^2}$

$$\frac{1}{e^{\beta \epsilon} - 1} \text{ vs } \frac{1}{2^{-1} e^{\beta \epsilon} - 1}$$

These 2 terms can differ appreciably

$$\downarrow z$$

$$1 + \beta \epsilon - z$$

$$z^2$$

We know $1-z$ is small but how small is it?

z may be $\sim \frac{1}{L}$

then we will ~~be~~ wrong in setting $z=1$

So we can't put $z=1$ in the 2nd term of the last page v.i.z., I_2

The limit is subtle
 $e(\vec{p}) \rightarrow 0$
 we don't know which is going faster

The no. of terms in the sum is unbounded as $L \rightarrow \infty$
 So it isn't enough to show that individual terms go to zero

$$\sum_{\vec{n} \neq 0} \frac{1}{n_1^2 + n_2^2 + n_3^2}$$

has vanishing contribution
as $L \rightarrow \infty$ due to
 \vec{n} factor

$$= \sum_{\substack{\vec{n} \neq 0 \\ |\vec{n}|^2 < M}} \frac{1}{n_1^2 + n_2^2 + n_3^2}$$

$M = A$ fixed large number

$$+ \sum_{\substack{M \leq |\vec{n}|^2 \leq \frac{m^2 \epsilon}{2\pi^2 h^2}}} \frac{1}{n_1^2 + n_2^2 + n_3^2}$$

$$\frac{1}{n^2} = \frac{m^2 \epsilon}{2\pi^2 h^2}$$

$$\int d^3 n \frac{1}{n^2}$$

$$= 4\pi \int \frac{\sqrt{m\epsilon}}{\sqrt{M}} \frac{L}{\sqrt{2\pi h}} d\vec{n} \cdot \frac{1}{n^2}$$

$$= 4\pi \left[\frac{L}{2\pi h} \sqrt{\frac{m\epsilon}{2}} - \sqrt{M} \right]$$

\therefore for large M ,
 $\frac{1}{n_1^2 + n_2^2 + n_3^2}$
 does not change
 appreciably on changing
 n_i by λ

(same limit as $\lim \vec{n} \rightarrow \infty$)

$$\lim_{\epsilon \rightarrow 0} \lim_{L \rightarrow \infty} I_2$$

$$= \lim_{\epsilon \rightarrow 0} \lim_{L \rightarrow \infty} \frac{m}{2\pi^2 h^2 \beta} \cdot \frac{1}{L} \left\{ 4\pi \left(\frac{L}{2\pi h} \sqrt{\frac{m\epsilon}{2}} - \sqrt{M} \right) + \text{finite} \right\}$$

$$= \lim_{\epsilon \rightarrow 0} \frac{m}{2\pi^2 h^2 \beta} \frac{4}{h} \sqrt{\frac{m\epsilon}{2}} = 0$$

(contribution which came from the sum $\sum_{|\vec{n}|^2 \geq M}$)

All the contribution must come from I_1

($T_b = 0$ term)

$$\frac{1}{\sqrt{2^{-1} - 1}} = \frac{4\pi}{c^3} k_B (2m\hbar)^{3/2} \left\{ T_c^{3/2} - T^{3/2} \right\}$$

$$\lim_{V \rightarrow \infty} \frac{1}{\sqrt{2^{-1} - 1}} =$$

$$\left(2^{-1} \approx 1 + \frac{c}{V} \right) \lim_{V \rightarrow \infty} \frac{1}{\sqrt{1 + \frac{c}{V} - 1}} = 1/c$$

The c limit
is not reached here

where $\frac{1}{c} = \frac{4\pi}{h^3} K_0 (2mk)^{3/2} \left[\frac{T_c^{3/2}}{T_c^{3/2} - T^{3/2}} \right]$

Fraction of particles at $\vec{p} = 0$ level

$$\frac{1}{N} \frac{1}{Z^{-1} - 1} = \frac{V}{N} \cdot \frac{1}{V} \frac{1}{Z^{-1} - 1}$$

$$-\frac{1}{\beta} \frac{4\pi}{h^3} K_0 (2mk)^{3/2} \left[\frac{T_c^{3/2}}{T_c^{3/2} - T^{3/2}} \right]$$

finite as $V \rightarrow \infty, N \rightarrow \infty$

$$\vartheta = \frac{N}{V} \text{ fixed}$$

\rightarrow Bose-Einstein condensation

$N_0 = \text{no. of particles in ground state}$
 $= \frac{z}{1-z}$
 $\Rightarrow z = \frac{N_0}{N_0 + 1} \rightsquigarrow \text{Pathria}$

Thermodynamics below T_c

$$Z^{-1} \approx 1 + c/N \quad Z \approx 1 - c/N$$

$$\begin{aligned} \ln \vartheta &= - \sum_{\vec{p}} \ln (1 - Z e^{-\beta E(\vec{p})}) \\ &= - \ln (1 - z) - \sum_{\vec{p} \neq 0} \ln (1 - z e^{-\beta E(\vec{p})}) \\ &\quad \rightarrow - \frac{V}{h^3} \int d^3 p \ln (1 - z e^{-\beta E_{2m}}) \\ &= \frac{4\pi V}{h^3} (2mkT)^{3/2} F_1(z) \end{aligned}$$

$\ln V$ becomes large in the thermodynamic limit but doesn't grow fast enough to affect the thermodynamics

$$F_1(z) = - \int_0^\infty u^2 du \ln (1 - z e^{-u^2})$$

The same treatment works for both $T < T_c$ & $T > T_c$ for $\ln \vartheta$ - no term contributes appreciably so that this is reqd.

so no special treatment reqd. for the $\vec{p} = 0$ term in this case
 this is due to the presence of \ln

$$P = kT \frac{\partial}{\partial V} \ln \Omega$$

$$= kT \cdot \frac{4\pi}{h^3} (2m\kappa T)^{3/2} f_1(z)$$

P , as it stands, doesn't depend on V .
 \sqrt{z} term has dropped out

Usually z is a fn. of $\frac{N}{V} \& T$ through the F_2 equation.
 This generates a dependence of P on $N/V \& T$

$$\text{For } T < T_c, z = 1 - c/N$$

$$F_1(z) \approx f_1(1) + G\left(\frac{1}{N}\right)$$

$$P = kT \cdot \frac{4\pi}{h^3} \cdot (2m\kappa T)^{3/2} f_1(1)$$

for $T < T_c$.

$$f_1(1) = - \int_0^\infty u^2 du \ln(1-e^{-u})$$

is a finite integral

→ independent of density

(this time it is truly indep. of density)

Strange bcs system bcs
 as you change vol. keeping
 the vol. fixed,
 $N \& T$ fixed,
 P doesn't change

$1 - e^{-Pc(\beta)}$
 $1 - e^{-Pc(\beta)}$
 changes or α charges

for $\alpha = 1$) the distribution
 is fixed. As you are compressing
 any more & more particles
 are accumulating in the
 zero energy state which
 don't exert pressure
 (density ↑)
 distribution of particles
 the exto states remain the
 same

This is not s.th. unusual, when we have a
 mixture of 2 phases (say, ~~the~~ liq. H_2O & vapour)
 - now reduce the vol. - pressure won't
 change, only ^{some} vapour will condense
 into liquid

$$\ln \Omega = \frac{4\pi V}{h^3} \left(\frac{2m}{\beta}\right)^{3/2} F_1(z)$$

$$E = \left(-\frac{\partial}{\partial \beta} \ln \Omega\right)_{Z,V}$$

$$= \frac{4\pi V}{h^3} \frac{3}{2\beta} \cdot \left(\frac{2m}{\beta}\right)^{3/2} F_1(z)$$

$= 1$
 for $T < T_c$ is a fn. of
 (N, β) for $T > T_c$

$$C_V = \left(\frac{\partial E}{\partial T} \right)_{N, V}$$

Ex. Check that for $T > T_c$

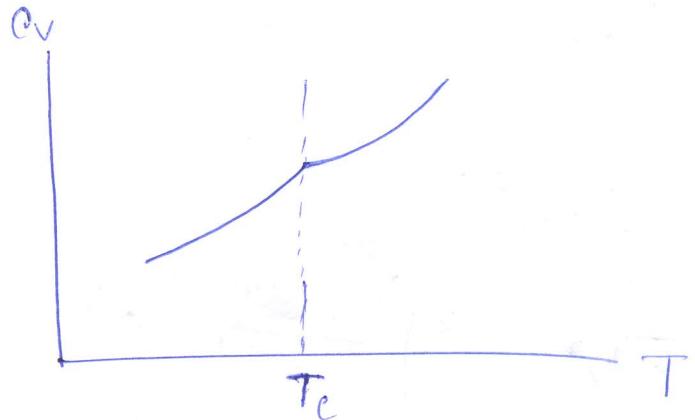
$$F_1(z) = F_1(1) + k_0(T - T_c)^2 + \dots$$

some constant

$$\frac{\partial}{\partial T} F_1(z) = C_0(T - T_c) \rightarrow 0 \text{ as } T \rightarrow T_c$$

(as a result) C_V is continuous at T_c .

$\frac{\partial C_V}{\partial T}$ is not continuous at T_c



C_V & all higher derivatives will have different behaviour for $T < T_c$ & $T > T_c$

H.W.

$$\begin{aligned}
 I &= \int_0^\infty dy y^2 \frac{1}{(e^{y\hbar} + e^{-y\hbar})^2} = \int_0^\infty dy \frac{y^2 e^y}{(1+e^y)^2} \\
 &= -\frac{y^2}{e^y + 1} \Big|_0^\infty + 2 \int_0^\infty \frac{y}{e^y + 1} dy \\
 &= 2 \int_0^\infty \frac{ye^{-y}}{1+e^{-y}} dy \\
 &= 2 \sum_{n=1}^\infty \int_0^\infty ye^{-ny} dy \\
 &= 2 \sum_{n=1}^\infty \left[-\frac{ye^{-ny}}{n} \Big|_0^\infty + \frac{1}{n} \int_0^\infty e^{-ny} dy \right] \\
 &= 2 \sum_{n=1}^\infty \left[\frac{1}{n^2} e^{-ny} \Big|_0^\infty \right] \\
 &= 2 \sum_{n=1}^\infty \frac{1}{n^2}
 \end{aligned}$$

$$\therefore \int_0^\infty dy y^2 \frac{1}{(e^{y\hbar} + e^{-y\hbar})^2} = 4 \sum_{n=1}^\infty \frac{1}{n^2} = \frac{4 \times \pi^2}{12} = \frac{\pi^2}{3}$$

H.W.
of 30/11/06

$$\begin{aligned}
 \frac{\sqrt{\pi}}{4} z \left(1 - \frac{z^2}{2\pi^2} + \dots\right) &= \frac{N}{V} \frac{h^3}{4\pi} (2m\hbar kT)^{-3/2} \\
 \Rightarrow \frac{\sqrt{\pi}}{4} z - \frac{\sqrt{\pi}}{8\sqrt{2}} z^2 &= \frac{N}{V} \frac{h^3}{4\pi} (2m\hbar kT)^{-3/2} \\
 \Rightarrow \frac{\sqrt{\pi}}{4} z &= \frac{\sqrt{\pi}}{8\sqrt{2}} z^2 + \frac{N}{V} \frac{h^3}{4\pi} (2m\hbar kT)^{-3/2} \\
 \Rightarrow \frac{\sqrt{\pi}}{4} z &= \frac{\sqrt{\pi}}{8\sqrt{2}} \times \left(\frac{N}{V}\right)^2 \left(\frac{2m\hbar kT}{h^2}\right)^{-3} \\
 &\quad + \frac{N}{V} \frac{h^3}{4\pi} (2m\hbar kT)^{-3/2}.
 \end{aligned}$$

for
successive
approximation
method

$$\begin{cases} \Rightarrow z = \frac{1}{2\sqrt{2}} \left(\frac{N}{V}\right)^2 \left(\frac{2m\hbar kT}{h^2}\right)^{-3} \\ \quad + \frac{N}{V} \frac{h^3}{4\pi\sqrt{2}} \left(\frac{h^2}{2m\hbar kT}\right)^{3/2} \end{cases}$$

$$Z = \frac{1}{2\sqrt{2}} \left(\frac{\bar{N}}{V} \right)^2 \left(\frac{h^2}{2m\pi kT} \right)^{3h} + \frac{\bar{N}}{V} \left(\frac{h^2}{2m\pi kT} \right)^{3h}$$

\therefore Eqn. of state is :-

$$\frac{PV}{\bar{N}kT} = 1 + \frac{1}{4\sqrt{2}} \frac{\bar{N}}{V} \left(\frac{h^2}{2m\pi kT} \right)^{3h}$$

$$+ \frac{1}{16} \left(\frac{\bar{N}}{V} \right)^2 \left(\frac{h^2}{2m\pi kT} \right)^{3h} + \dots$$

$$\frac{4\pi^2 k^2}{2m\pi kT} = \frac{2\pi k^2}{m\pi kT}$$

⊗ $\frac{4\pi^2}{4\sqrt{2}}$

~~10/2/06~~

System of photons in a box

Quanta of electromagnetic radiation

- ① Photons are massless particles

$$E(\vec{p}) = \sqrt{m^2 c^4 + \vec{p}^2 c^2} = c |\vec{p}|$$

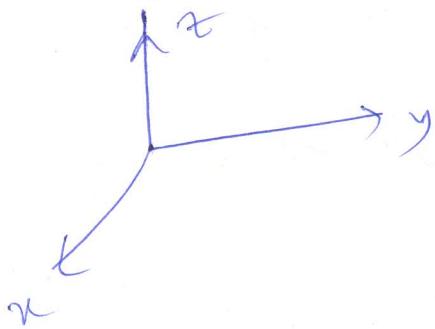
- ② Each photon comes in two polarization states.

Need another quantum number besides momentum to specify the state of a photon.

$s = +1$ or -1
 \rightarrow represent the result of quantizing left circularly polarized and right circularly polarized em wave.

$$(E_x, B_y) \neq 0$$

$$(E_y, B_x) \neq 0$$



$$\sum_{\vec{p}} \rightarrow \sum_{s=\pm 1} \sum_{\vec{p}}$$

- ③ Number of photons is not conserved.
 \rightarrow can get emitted and absorbed by the wall.
 \rightarrow cannot work with a fixed no. of photons.

Recall standard procedure for grand canonical ensemble:

$$Q = \sum_N T_N (e^{-\beta H + \mu N})$$

Determine μ by requiring that

$$\sum \text{Tr} (e^{-\beta H + \beta \mu n})$$
$$\bar{N} = \frac{\sum_N N \text{Tr} (e^{-\beta H + \beta \mu n})}{\sum_N \text{Tr} (e^{-\beta H + \beta \mu n})}$$

μ was a parameter we introduced by hand, but ultimately

μ was eliminated or μ had to be determined by demanding $\bar{N} = \dots$

by solving that eqn.

agrees with total no. of particles

This procedure will not work bcos u can't fix \bar{N} for photons
- we can't determine μ

Recall the canonical partition fn:

$$Z(\beta) = \text{Tr} (e^{-\beta H})$$

↓ over all states with arbitrary number of photons

Previously, we restricted the Trac to be taken over a fixed no. of particles - but now it's not true

$$= \mathcal{G} (\mu = 0)$$

this implies, for photons we can use grand canonical part'n function if we set $\mu = 0$.

In case of photons, we don't try to fix \bar{N} . Here μ is our input parameter & that is equal to zero

This is true for all systems where no. of particles of a particular type are conserved

It's a generic feature for all systems for which the number of particles is not conserved,

In a system where $A+B \rightarrow C+D$
Break up the cond. charges - those that are conserved are det. by $\bar{N} = \dots$
& those that aren't, we set $\mu \neq 0$

N is still the no. of particles but it no longer is the input parameter
— we don't try to match N with the total no. of particles in the system

$$\ln Q = - \sum_{s=+1,-1} \sum_{p \in \infty} \ln (1 - e^{-\beta e(p)})$$

$$= -2 \frac{4\pi V}{h^3} \int_0^\infty p^2 dp \ln (1 - e^{-\beta e(p)})$$

(Here we are in a dangerous situation as by putting $\mu = 0$, we have set $Z = 1$)

— But still this integral can be written bcos in $\log Q$ there is no problem)

~~degeneracy~~

$$\beta c p = u$$

$$\text{But } e(p) = cp \quad p = \frac{u + T}{c}$$

$$\ln Q = \frac{8\pi V}{h^3} \cdot \left(\frac{kt}{c}\right)^3 \left[- \int u^2 du \ln (1 - e^{-u}) \right] A \quad (\text{some number})$$

$$\frac{\partial}{\partial T} \ln Q = \frac{\partial}{\partial V} \ln Q = \frac{8\pi}{h^3} \left(\frac{kt}{c}\right)^3 A$$

$$F = - \frac{\partial \ln Q}{\partial \beta} = kT^2 \frac{\partial}{\partial T} \ln Q$$

$$= \frac{24\pi V}{h^3 c^3} A (kt)^4$$

$$f = \frac{F}{V} = \frac{24\pi}{h^3 c^3} (kt)^4 \cdot A$$

$\int d^4 T^4$ Stefan-Boltzmann law

$\phi = 8/3 \rightarrow$ known relation b/w pressure & energy density of radiation.

pressure doesn't depend upon volume
— this isn't surprise bcos in a normal gas, as $V \downarrow$ (T fixed) $N \uparrow$ (density \uparrow)
— but here as $V \downarrow$, no. of photons \downarrow appropriately as density of photons \uparrow (intrinsic prop. of the system)

Any relation which doesn't have an "i" is a classical relation

avg. occupancy no. of a mom. state per polarization in the thermodynamic limit, it is not reasonable to focus on a single state or a few particles in the mom. range $\langle n \rangle = \frac{1}{h^3} \int_{-h/2}^{h/2} \int_{-h/2}^{h/2} \int_{-h/2}^{h/2} \delta(\vec{p} - \vec{p}_0) \delta(\vec{p} - \vec{p}_1 + \vec{p}_2)$

$$n(\vec{p}, \beta) = \frac{e^{-\beta E(\vec{p})}}{1 - e^{-\beta E(\vec{p})}} = \frac{e^{-\beta c p}}{1 - e^{-\beta c p}}$$

Total no. of states in momentum range $p \leq |\vec{p}| \leq p + \delta p = \frac{8\pi V p^2 \delta p}{h^3}$

$$\sum_s \sum_{\vec{p}} \rightarrow \int \frac{8\pi V p^2 dp}{h^3}$$

\Rightarrow Average no. of photons with $p \leq |\vec{p}| \leq p + \delta p$

$$\langle n \rangle = \frac{8\pi V p^2 \delta p}{h^3} \frac{e^{-\beta c p}}{1 - e^{-\beta c p}}$$

Average energy carried by photons in the momentum range $p \leq |\vec{p}| \leq p + \delta p$

$$= \frac{8\pi V}{h^3} p^2 \delta p \frac{e^{-\beta c p}}{1 - e^{-\beta c p}} cp$$

Multiply
avg. # of photons
--- with cp

Quantum mechanics : (relates)

$$E = h\nu \rightarrow \text{frequency of the wave}$$

energy of photon

one side q. mech. & on the other side class. em

$$p = h/\lambda \rightarrow \text{wave-length}$$

momentum

$$E = pc \Rightarrow \nu = c/\lambda$$

Sum over s
 s will be non-trivial
if you somehow remove the degeneracy & make it very
e.g. get charge in electric field

Average energy in photons in the wavelength range $(\lambda, \lambda + \delta\lambda)$

$$= \frac{8\pi V}{\lambda^3} \left(\frac{h}{\lambda}\right)^2 \hbar \left(\frac{\delta\lambda}{\lambda^2}\right) e^{-\frac{ch}{kT}} \frac{e^{-\frac{ch}{kT+\delta\lambda}}}{1 - e^{-\frac{ch}{kT+\delta\lambda}}} \times c \frac{h}{\lambda}$$

$$\boxed{p \rightarrow h\nu \\ \delta\lambda \rightarrow \frac{h}{\lambda^2} \delta\lambda}$$

$$= \frac{8\pi V \cdot ch}{\lambda^5} \underbrace{\frac{e^{-\frac{ch}{kT}}}{1 - e^{-\frac{ch}{kT}}}}_{f(\lambda)}$$

\rightarrow Planck distribution law

- 'sign not imp. as you are measuring the magnitude'

High temperature limit: $T \rightarrow \infty$

$$f(\lambda) \rightarrow \frac{8\pi ch}{\lambda^3} \frac{1}{e^{\frac{ch}{kT}}} = \frac{8\pi kT}{\lambda^4}$$

This formula doesn't have an 'h' & can be derived from class stat. mech. for em fields - there is a classical origin of the formula

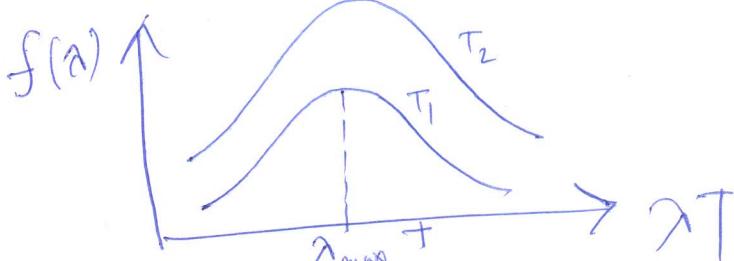
No. of modes of an em field \propto can change - valid in classical picture - but from the pt. of view of particles, how to sum over no. of photons not conserved. so here high temp. limit is class stat. mech. of fields & not particles can't draw analogies not $\propto h \rightarrow$ give quant. theory should give vs classical picture

Full formula for f :

$$f(\lambda) = \frac{8\pi ch}{\lambda^5} \frac{e^{-\frac{ch}{kT}}}{1 - e^{-\frac{ch}{kT}}}$$

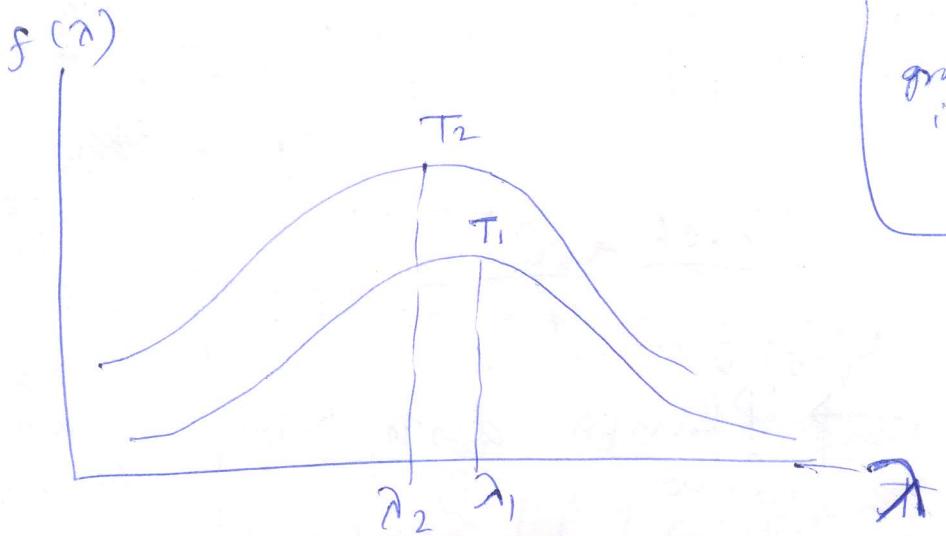
$$= T^5 \cdot 8\pi ch \left\{ \frac{e^{-\frac{ch}{kT}}}{(AT)^5 (1 - e^{-\frac{ch}{kT}})} \right\}$$

or $f(\lambda) \rightarrow$ a fn. of AT



$$\lambda_{max}^2 \quad T_1 = \lambda_{max}^2 T_2$$

$f(\lambda) = T^5 \downarrow$ such $g(\lambda T) \rightarrow$ so peak of
 overall multiplicative factor f vs λT at same value of λT



graphs are identical upto a scaling

$$\lambda_1 T_1 = \lambda_2 T_2$$

$$f(\lambda) = T^5 \text{ such } g(\lambda T)$$

$$\frac{\partial f}{\partial \lambda} = 0 \Rightarrow g'(\lambda T) = 0$$

$g'(u) = 0$ has a soln. $u = u_0$
ch/ku
 fixed no.

$$g(u) = \frac{1}{u^5} \frac{e^{-\frac{1}{u}}}{1 - e^{-\frac{1}{u}}}$$

$$\lambda_{\max} \cdot T = u_0$$

→ position of maximum

Bose-Einstein condensation isn't seen in photons - infinite no. of photons in $\epsilon=0$ state has no physical effect - you may ask whether there is any significant effect from levels close to zero but not zero energy - this will give negligible effect as in the case we dealt with

H Electrons in a magnetic field

Ignore spin
Uniform magne'tic field ~~along z-dirn~~

$$B_x = B_y = 0, B_z = B$$

$$\vec{B} = \vec{\nabla} \times \vec{A}$$

Choose $A_x = -\frac{1}{2}B_y, A_y = \frac{1}{2}B_x, A_z = 0$

$$\Rightarrow B_z = \partial_x A_y - \partial_y A_x = B$$

$\partial_x = \frac{\partial}{\partial x}$

Check $B_x = 0, B_y = 0$

Consider an electron of charge $-e$
moving in this field

$$H = \frac{1}{2m} (\vec{p} + e\vec{A}/c)^2$$

$$= \frac{1}{2m} \left\{ p_x^2 + \left(p_x - \frac{1}{c} e B_y \right)^2 + \left(p_y + \frac{1}{c} e B_x \right)^2 \right\}$$

$$= \underbrace{\frac{p_z^2}{2m}}_{H_z} + \underbrace{\frac{1}{2m} \left[p_x^2 + p_y^2 + \frac{e^2 B^2}{4c^2} (x^2 + y^2) \right]}_{H_0} + \underbrace{\frac{eB}{2mc} (xp_y - yp_x)}_{L_z}$$

H_z, H_0 & L_z commute with each other

\Rightarrow can be simultaneously diagonalized

$\omega = \frac{eB}{2mc}$

 $\rightarrow \text{defn} \rightarrow$

$K = \frac{e^2 B^2}{4mc^2}$

Creation &
annihilation
op'

$$+\vec{a}_x = \frac{1}{\sqrt{2}} \left\{ \left(\frac{h}{km} \right)^{1/2} p_x + i \left(\frac{km}{h} \right)^{1/2} \alpha_x \right\}$$

$$-\vec{a}_x = \frac{1}{\sqrt{2}} \left\{ \left(\frac{h}{km} \right)^{1/2} p_x - i \left(\frac{km}{h} \right)^{1/2} \alpha_x \right\}$$

a_y, a_y^+ similarly

$$\begin{aligned} \hat{a}_x^{\dagger} a_y &= \frac{1}{2} \left[\frac{t}{\hbar m} p_x p_y + \left(\frac{\hbar m}{t} \right) \omega y + i \omega p_y \right] \\ &= \frac{1}{2} \left[\left(\frac{\hbar m}{t} \right) p_x p_y + \left(\frac{\hbar m}{t} \right) \omega y - L_z / \hbar \right] \\ a_y^{\dagger} a_x &= \frac{1}{2} \left[\left(\frac{\hbar m}{t} \right) p_x p_y + \left(\frac{\hbar m}{t} \right) \omega y + L_z / \hbar \right] \\ \therefore a_x^{\dagger} a_y - a_y^{\dagger} a_x &= -L_z / \hbar \Rightarrow L_z |0\rangle = 0 \end{aligned}$$

$$\mathcal{H}_0 = \hbar \omega (a_x^{\dagger} a_x + a_y^{\dagger} a_y + 1)$$

$$a_x |0\rangle = 0 = a_y |0\rangle \Rightarrow L_z |0\rangle = 0 \quad (\text{Check this})$$

Ground state
Excited states

$(a_x^{\dagger})^{n_1} (a_y^{\dagger})^{n_2} |0\rangle$ eigenstates
→ not L_z

$$[L_z, a_n^+] = i\hbar a_n^+, [L_z, a_n^+] = -i\hbar a_n^+$$

$$\text{Define: } a_{+}^{\dagger} = \frac{1}{\sqrt{2}} (a_x^{\dagger} + i a_y^{\dagger})$$

$$a_{-}^{\dagger} = \frac{1}{\sqrt{2}} (a_x^{\dagger} - i a_y^{\dagger})$$

Defining new creation & annihilation operators

These have the following features:-

$$\textcircled{1} \quad [a_{+}, a_{+}^{\dagger}] = 1, [a_{-}, a_{-}^{\dagger}] = 1$$

other commutators vanish

$$\textcircled{2} \quad [L_z, a_{+}^{\dagger}] = \hbar a_{+}^{\dagger}$$

$$[L_z, a_{-}^{\dagger}] = -\hbar a_{-}^{\dagger}$$

$$\textcircled{3} \quad \mathcal{H}_0 = \hbar \omega (a_{+}^{\dagger} a_{+} + a_{-}^{\dagger} a_{-} + 1)$$

$$\begin{aligned} H_{\text{new}}: L_z a_{+}^{\dagger} |0\rangle - a_{+}^{\dagger} L_z |0\rangle &= \hbar a_{+}^{\dagger} |0\rangle \\ \Rightarrow L_z a_{+}^{\dagger} |0\rangle &= \hbar a_{+}^{\dagger} |0\rangle \\ &\text{Similarly,} \\ L_z a_{-}^{\dagger} |0\rangle &= -\hbar a_{-}^{\dagger} |0\rangle \end{aligned}$$

Define $|0\rangle$ such that $a_{+}^{\dagger} |0\rangle = 0 = a_{-}^{\dagger} |0\rangle$

$$(a_{+}^{\dagger})^{N-n_1} (a_{-}^{\dagger})^{n_1} |0\rangle = |s\rangle$$

$$\begin{aligned} \mathcal{H}_0 |s\rangle &= \hbar \omega (N - n_1 + n_1 + 1) |s\rangle \\ &= (N+1) \hbar \omega |s\rangle \end{aligned}$$

$$\begin{aligned} L_z |s\rangle &= \hbar \{ (N - n_1) - n_1 \} |s\rangle \\ &= \hbar (N - 2n_1) |s\rangle \end{aligned}$$

\therefore Lz has eigenvalues with eigenstates $\hbar \omega (N - n_1 + n_1 + 1)$

$N = 1, 2, 3, \dots$
 $n_1 = 0, 1, 2, \dots, N$

~~(H₀ + L_z) | 1s > | 2s >~~

~~(H₀ + eB/2mc L_z) | 1s > | 2p >~~

$$\left(H_0 + \frac{eB}{2mc} L_z \right) = \left(H_0 + \omega L_z \right)$$

$$\begin{aligned} L_z (a_{+}^{\dagger})^2 | 0 \rangle &= 2\hbar (a_{+}^{\dagger})^2 | 0 \rangle \\ L_z (a_{+}^{\dagger})^3 | 0 \rangle &= 3\hbar (a_{+}^{\dagger})^3 | 0 \rangle \\ L_z (a_{+}^{\dagger})^n | 0 \rangle &= n \hbar (a_{+}^{\dagger})^n | 0 \rangle \end{aligned}$$

~~(H₀ + ω L_z) | 1s >~~

$$\begin{aligned} &= \hbar \omega (N + 1 + (N - 2n_1)) | 1s \rangle \\ &= \hbar \omega (2N - 2n_1 + 1) | 1s \rangle \\ &\quad \underbrace{(2j+1)}_{(2j+1)} \end{aligned}$$

Eigenvalues of $\left(H_0 + \frac{eB}{2mc} L_z \right)$ are
 $(2j+1) \omega \hbar$, $j = \text{integer.}$

$$N - n_1 = j$$

j runs from 0 to N

Every j can be obtained in infinite no. of ways.

H.W. $x = \sqrt{\frac{\hbar}{2m\omega}} (a + a^{\dagger})$ & $p = -i\sqrt{\frac{m\omega}{2}} (a^{\dagger} - a)$

$$x p_y - y p_x$$

$$= i\hbar L_z [(a_y + a_y^{\dagger})(a_x^{\dagger} - a_x) - (a_x a_x^{\dagger} - a_x^{\dagger} a_x)(a_y^{\dagger} - a_y)]$$

$$= i\hbar L_z [a_y a_x^{\dagger} - a_y a_x^{\dagger} + a_y^{\dagger} a_x^{\dagger} - a_y^{\dagger} a_x - a_x a_y^{\dagger} + a_x a_y - a_x^{\dagger} a_y^{\dagger} + a_x^{\dagger} a_y]$$

$$= i\hbar (a_x^{\dagger} a_y - a_y^{\dagger} a_x)$$

$$\therefore L_z | 0 \rangle = 0 \quad (\because a_y | 0 \rangle = a_x | 0 \rangle = 0)$$

a_x & a_y commute

similarly,
 a_x^{\dagger} & a_y^{\dagger} commute

$$(L_z, a_x^{\dagger}) = i\hbar a_y^{\dagger} \Rightarrow L_z a_x^{\dagger} - a_x^{\dagger} L_z = i\hbar a_y^{\dagger}$$

$$\Rightarrow L_z a_x^{\dagger} | 0 \rangle - a_x^{\dagger} L_z | 0 \rangle = i\hbar a_y^{\dagger} | 0 \rangle \Rightarrow L_z a_x^{\dagger} | 0 \rangle = i\hbar a_y^{\dagger} | 0 \rangle$$

Similarly, $L_z a_y^{\dagger} | 0 \rangle = i\hbar a_x^{\dagger} | 0 \rangle$