

Comparison between 10-d and 4-d viewpoints

~~Only~~ Assume that  $\rho$  is the only  
Kähler moduli ( $h_{1,1}=1$ )

Results in 4-d:

$$V = e^{+k} (G^{i\bar{j}} \partial_i D_{\bar{j}} W \overline{D_{\bar{j}} W} - 3|W|^2)$$

$$D_i W = \partial_i W + \partial_i k W.$$

$$k = -3 \ln(-i(\rho - \bar{\rho})) = \ln(-i(\tau - \bar{\tau}))$$

$$\phi = \ln \left( -\frac{i}{2} \int \Omega \wedge \bar{\Omega} \right)$$

$$W = \int G^{(3)} \wedge \Omega, \quad G^{(3)} = F^{(3)} - \tau H^{(3)}$$

$$i: \rho, \alpha$$

↳ complex structure,  $\alpha$ .

$$D_{\rho} W = \partial_{\rho} W - \frac{3}{\rho - \bar{\rho}} W = -\frac{3}{\rho - \bar{\rho}} W$$

Substitute into  $V$ :

$$V = e^{+k} G^{\alpha\bar{\beta}} D_{\alpha} W \overline{D_{\alpha} W}$$

→ independent of  $\rho$ .

Minimization:  $D_{\alpha} W = 0$ .



Unbroken SUSY:

$$D_\mu W = 0 \quad \forall \mu \rightarrow \text{general result from SUGRA.}$$

$$D_\mu W = 0 \Rightarrow W = 0.$$

$$\Rightarrow \text{SUSY condition: } W = 0, D_\alpha W = 0.$$

Now compare with results from 10-d:

① We have already assumed that we have solved for background  $\tau_{10-d}$  and metric on  $CY_3$ . ( $F^{(3)}, H^{(3)}$  do not enter explicitly)

Source of Einstein eq. involves:

$$\hat{T}_{mn}^{\text{loc}} - \frac{1}{8} \hat{g}_{mn} (\hat{T}^{\text{loc}})^m_m$$

Vanish for 3-brane (only  $\hat{T}_{\mu\nu} \neq 0$ ).

② In this background switch on  $F^{(3)}, H^{(3)}$ .

$$\text{Eqs. of motion} \Rightarrow *G^{(3)} = i G^{(3)}$$

$G^{(3)}$  is imaginarily anti-self-dual.



Vanishing of total D3-brane charge

$$\int d^4y \sqrt{\det \tilde{g}} \left[ \frac{1}{12\kappa_{10}^2} \tilde{g}^{mn} \tilde{g}^{pq} \tilde{g}^{rs} \tilde{g}^{t\bar{t}} \right. \\ \left. G_{mnp} \star \bar{G}_{m'p'q'} + 2\kappa_{10}^2 e^{-6A} T_3 \phi_3^{\text{loc}} \right]$$

satisfy  $\Downarrow$  by plane localized D3-branes  
(03 planes).

$\sim$  will not affect the metric  $\tilde{g}_{mn}$ .

Finally solve eq. for  $\alpha = e^{4A}$   
 $\swarrow$  from  $F^{(5)}$   $\searrow$  Warp factor of the metric

$$ds^2 = e^{2A(y)} \eta_{\mu\nu} dx^\mu dx^\nu + e^{-2A} \tilde{g}_{mn}(y) dy^m dy^n$$

/ physical.

Some mathematical results on CY<sub>3</sub>:

①  $\star G = i G$

$\Rightarrow G$  is a linear combination of  $\binom{0,3}{\text{3,0}}$  and  $\binom{2,1}{\text{1,2}}$  forms.

② If  $\Omega$  is a (3,0) form then  $\partial \Omega$  is an l.c. of (3,0) and (2,1) forms.

complex structure



③ Unbroken SUSY

$G^{(3)}$  is a ~~(3,0)~~  $\binom{2,1}{\text{form}}$ .

(no  $\binom{0,3}{\text{component}}$ ).

Examine

$$W = \int G^{(3)} \wedge \Omega$$

$$\partial_A W = \int G^{(3)} \wedge \partial_A \Omega$$

$$\partial_A K = - \ln \left( -\frac{i}{2} \int \Omega \wedge \bar{\Omega} \right)$$

$$\partial_A K = - \frac{\int \partial_A \Omega \wedge \bar{\Omega}}{\int \Omega \wedge \bar{\Omega}}$$

$$\Rightarrow \mathcal{D}_A W = \partial_A W + \partial_A K W$$

$$= \int G^{(3)} \wedge \partial_A \Omega - \frac{\int \partial_A \Omega \wedge \bar{\Omega}}{\int \Omega \wedge \bar{\Omega}} \int G^{(3)} \wedge \Omega$$

$G^{(3)}$  has  $\binom{0,3}{\text{component}}$  &  $\binom{2,1}{\text{component}}$ .

$\partial_A \Omega$  has  $(3,0)$  and  $(2,1)$  component.

$\Rightarrow$  Only the  $\binom{0,3}{\text{component}}$  of  $G^{(3)}$  contributes

~~(3,0)~~  $\rightarrow$   $c \bar{\Omega}$

$c \bar{\Omega}$

some constant.

$$\begin{aligned} \mathcal{D}_A W &= c \int \bar{\Omega} \wedge \partial_A \Omega - \frac{\int \partial_A \Omega \wedge \bar{\Omega}}{\int \Omega \wedge \bar{\Omega}} c \int \bar{\Omega} \wedge \Omega \\ &= c \int \bar{\Omega} \wedge \partial_A \Omega - c \int \bar{\Omega} \wedge \partial_A \Omega = 0. \end{aligned}$$



$$A^{(3)} = F^{(3)} - \tau H^{(3)}$$

$$\begin{aligned} \partial_\tau W &= \int \partial_\tau A^{(3)} \wedge \Omega \\ &= - \int H^{(3)} \wedge \Omega \end{aligned}$$

$$D_\tau W = \partial_\tau W + \partial_\tau k W$$

$$= - \int H^{(3)} \wedge \Omega - \frac{1}{\tau - \bar{\tau}} \int A^{(3)} \wedge \Omega$$

$$= - \int \frac{(\bar{A}^{(3)} - A^{(3)})}{\tau - \bar{\tau}} \wedge \Omega - \frac{1}{\tau - \bar{\tau}} \int A^{(3)} \wedge \Omega$$

$$= - \frac{1}{\tau - \bar{\tau}} \int \bar{A}^{(3)} \wedge \Omega = 0$$

$$\checkmark \quad \text{Lo } (3,0)$$

$$\otimes (3,0) + (1,2)$$

Thus eq. of motion from 10-d perspective  $\Rightarrow$  4-d eq. of motion.

SUSY

$A^{(3)}$  has only (2,1) component.

$$\Rightarrow W = \int A^{(3)} \wedge \Omega = 0$$

$\rightarrow$  agrees with 4-d result.



So far we have fixed complex structure moduli  $\tau$ .

~~What is the order of the masses?~~

~~$x = e^{\tau}$  is independent term.~~

~~$\frac{1}{2} \ln \frac{3}{4} \ln \frac{4}{3}$~~

~~Ans.~~

For having control we need  $\tau_2$  large.

— Not true in general.

But for appropriate choice of fluxes  $\tau_2$  may be moderately large.

Invoke this mechanism.

$e_0$  is still not determined.

Not possible using perturbative means.

KKLT: Use non-perturbative effects.

LVS: Use  $\alpha'$  corrections.



KKLT

Non-perturbatively generated superpotential

← e.g. Euclidean D3 on 4-cycle of  $CY_3$ .

$$\text{Area} \propto R^4 \propto \text{Im} \rho$$

$$\Delta W \propto e^{i c \rho}$$

constant.

Look for <sup>susy</sup> solns. to  $D_\alpha W = 0$ ,  $D_{\bar{\alpha}} W = 0$   
^  
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Zereth order soln. exists.

$\Delta W$  corrects it.

Solve & 'integrate out' the  $\tau$   
and complex structure moduli.

$$\Rightarrow W_{\text{eff}} = \underbrace{W_0}_{\text{Constant}} + \underbrace{A}_{\text{Constant}} e^{i c \rho}$$

$$K = -3 \ln(-i(c\rho - \bar{\rho}))$$



$$D_\rho W_{\text{eff}} = 0$$

$$\Rightarrow \partial_\rho W_{\text{eff}} + \partial_\rho K W_{\text{eff}} = 0.$$

$$icA e^{ic\rho} - \frac{3}{e-\bar{e}} (W_0 + A e^{ic\rho}) = 0.$$

Take  $c, A, W_0$  real.

Look for soln.  $\rho = i\sigma$   
 $\sigma$  real.

$$\Rightarrow cA e^{-c\sigma} + \frac{3}{2\sigma} (W_0 + A e^{-c\sigma}) = 0$$

$$\Rightarrow W_0 = -\frac{2\sigma}{3} cA e^{-c\sigma} - A e^{-c\sigma}$$

$$= -A e^{-c\sigma} \left( 1 + \frac{2c\sigma}{3} \right)$$

$$V = e^K (|D_\rho W|^2 - 3|W|^2)$$

$$= -\frac{c^2 A^2 e^{-2c\sigma}}{6} \Big|_{\sigma = \sigma_{\text{cr}} \text{ soln.}}$$

$\rightarrow$  AdS vacuum.



Now we ~~have~~ have fixed all moduli.

→ need to lift it to dS.

Strategy: ~~B~~ Adjust fluxes so that we have large warping.

~~A (E)~~ Placing D3-brane at places of small  $g_{\text{uv}}(y) \Rightarrow$  small tension.  
Tautg small.

Place D3-brane there to cancel the negative  $V$  & produce small +ve  $V$ .

(Adjust fluxes to satisfy D3-brane charge = 0 condition.)